

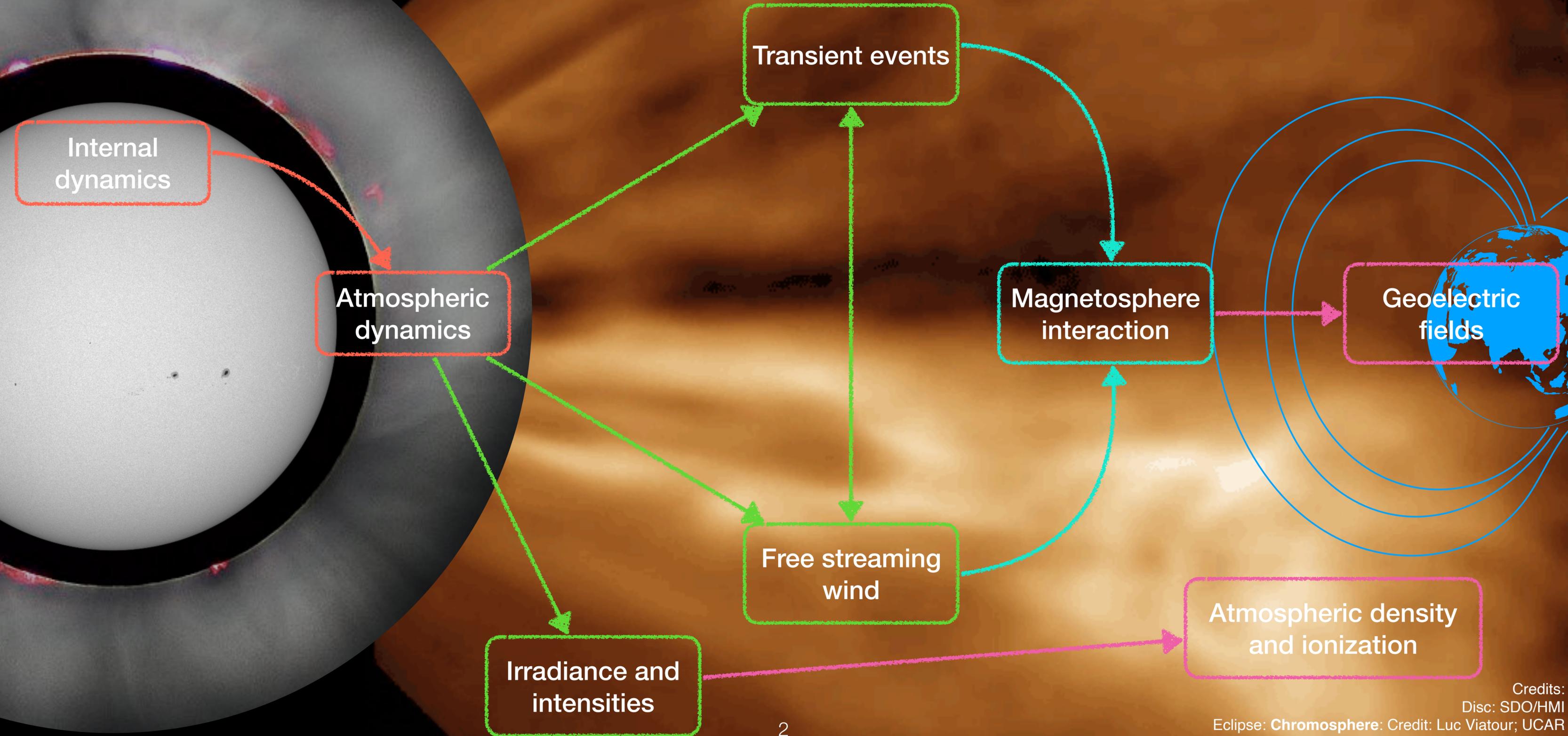
vishal@lmsal.com

From subtle to the vast: A multi-scale understanding of the Sun aided by artificial intelligence

Vishal Upendran

Lockheed Martin Solar and Astrophysics Laboratory | [Bay Area Environmental Research Institute](#)

Heliosphere as a system [Certainly NOT to scale!]



AI applications to Heliosphere

Simulation-based inference

Explainable AI for model evaluation and scientific discovery

Physics-incorporated deep learning

Atmospheric dynamics

Free streaming wind

Magnetosphere interaction

Length: $\sim 10^6$ m, Time: \sim seconds to mins

Simulation-based inference using machine learning

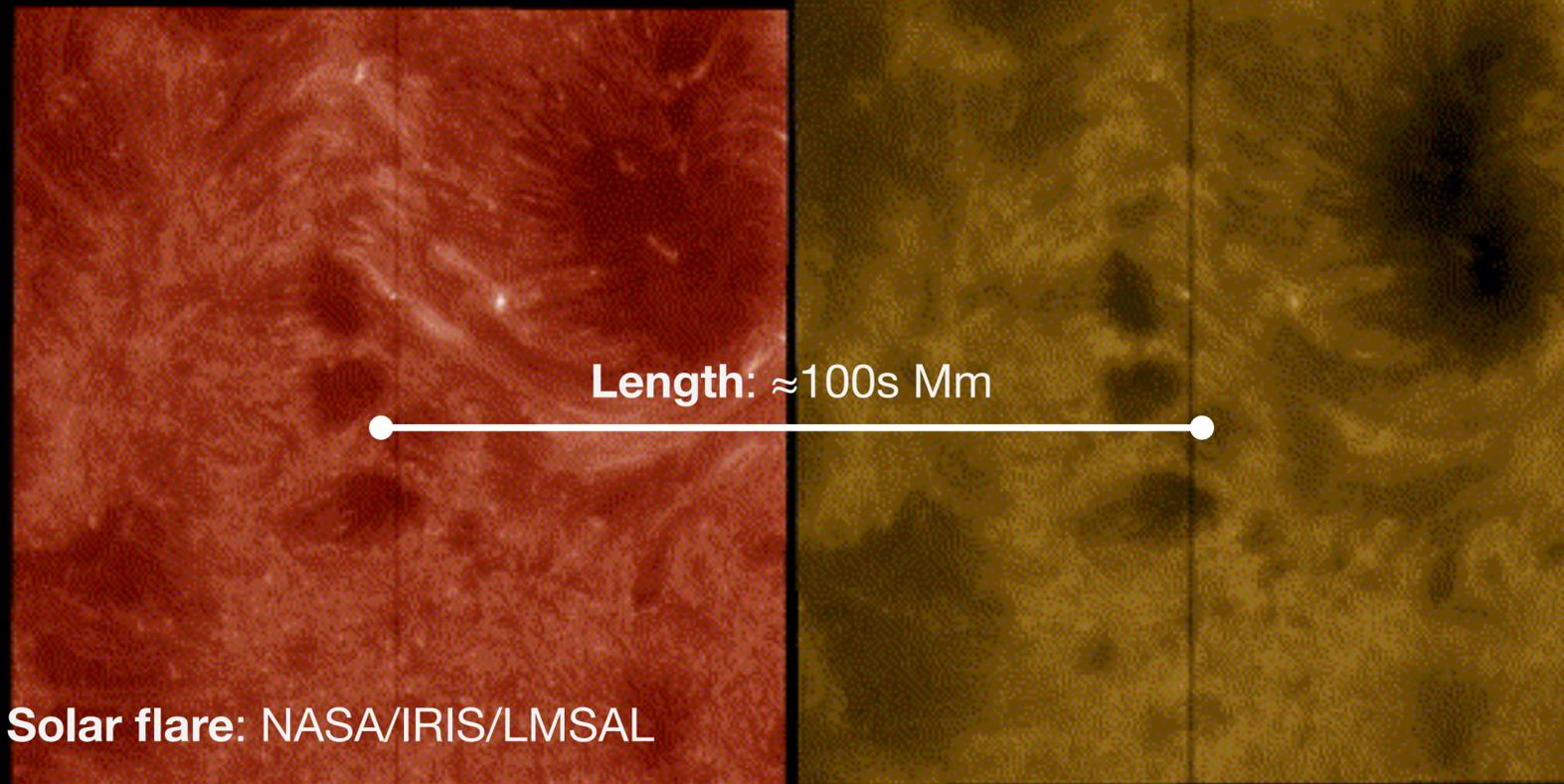
Application: Parameter estimation for $>1,000,000$ light curves

Upendran and Tripathi
(2021) ApJ 916 59

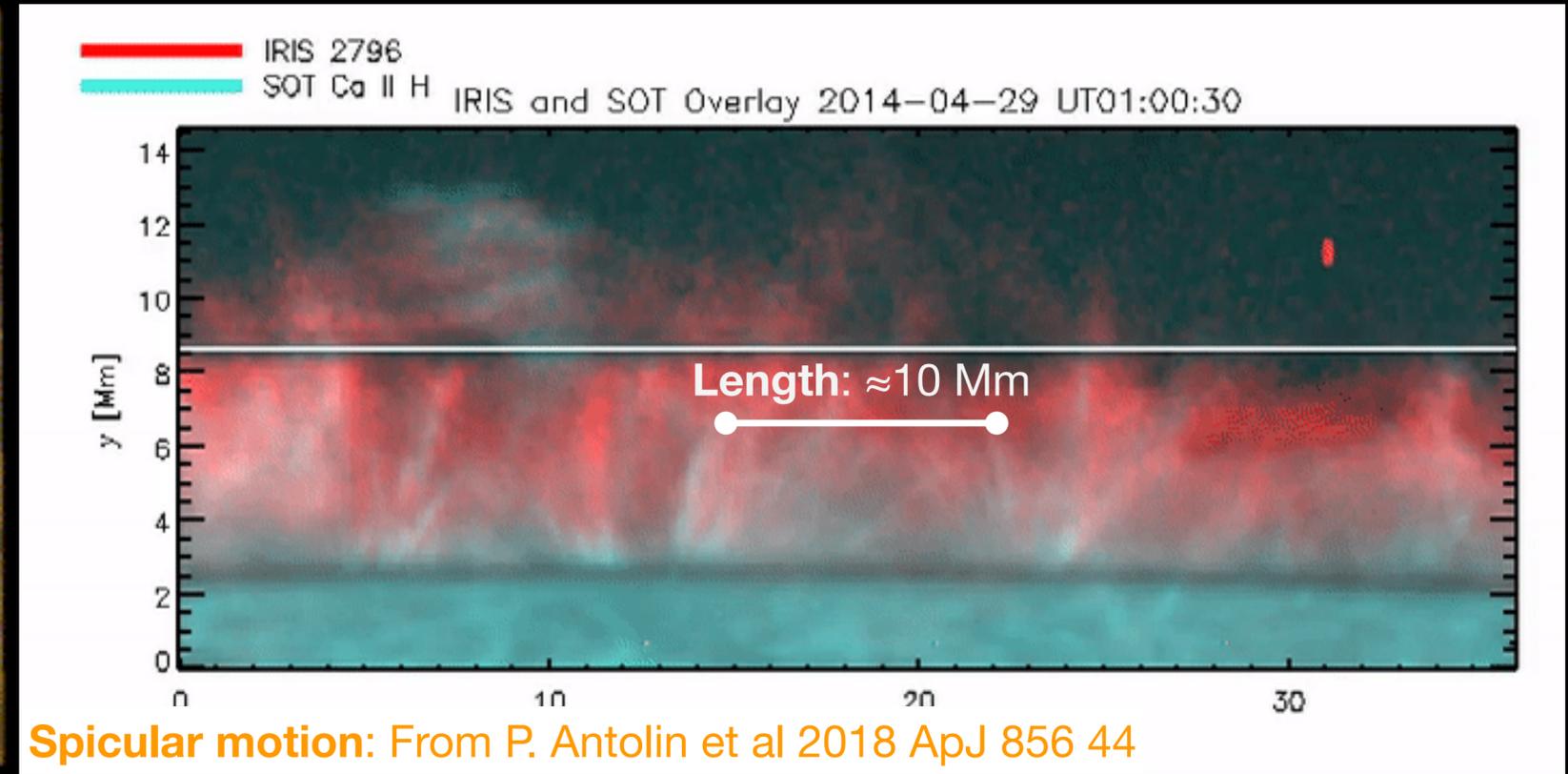


Upendran et. al. (2022)
ApJL 940 L38

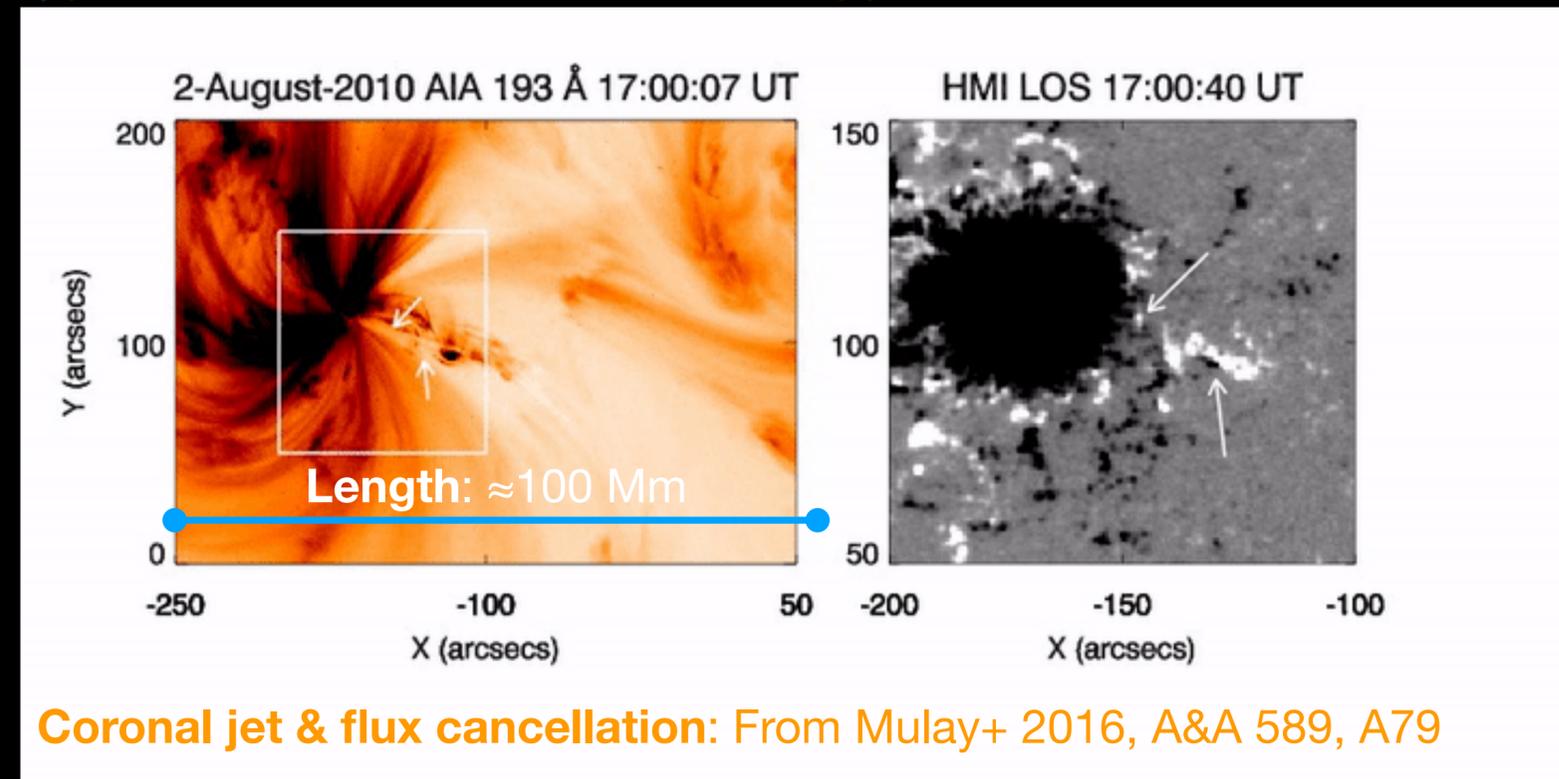
Energy release at multiple scales



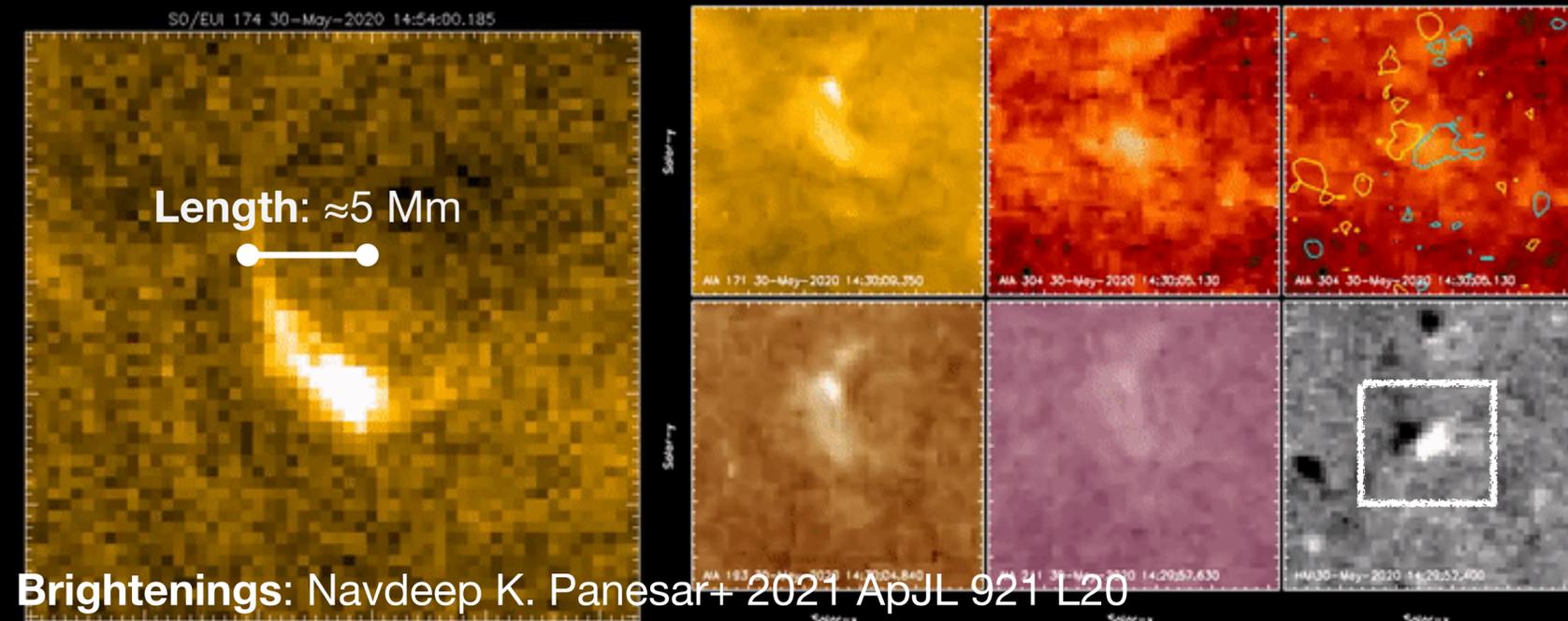
Solar flare: NASA/IRIS/LMSAL



Spicular motion: From P. Antolin et al 2018 ApJ 856 44



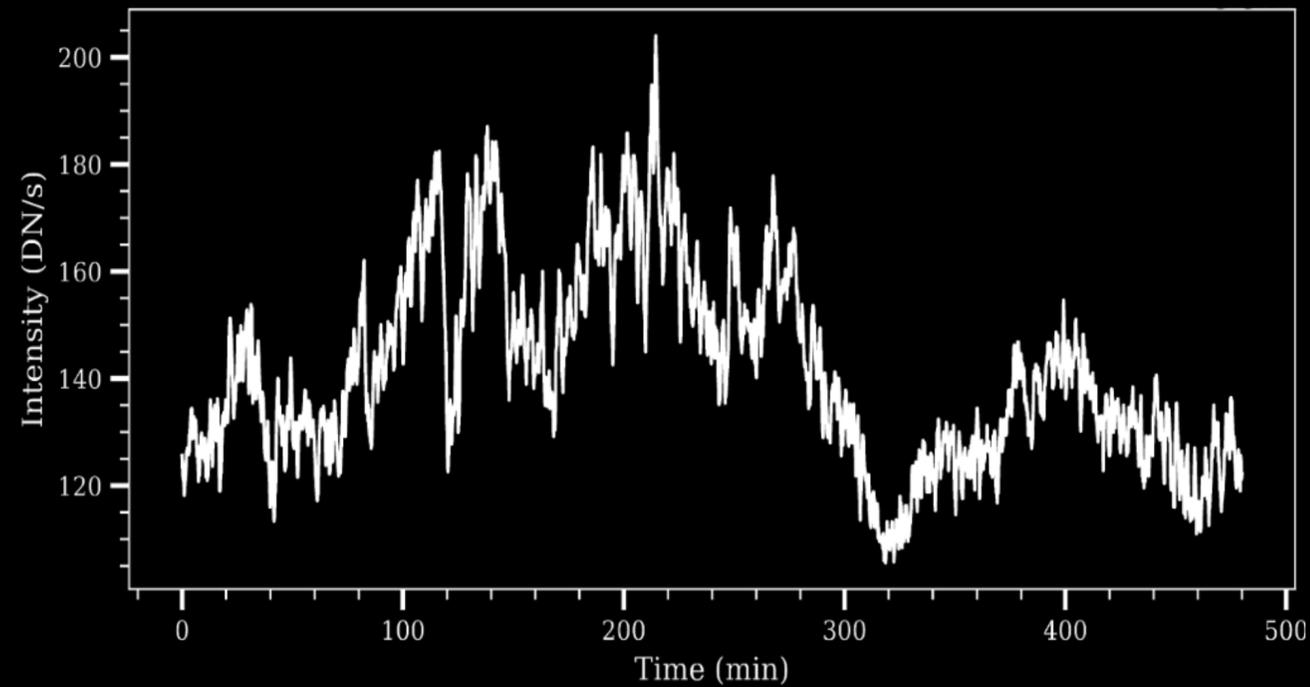
Coronal jet & flux cancellation: From Mulay+ 2016, A&A 589, A79



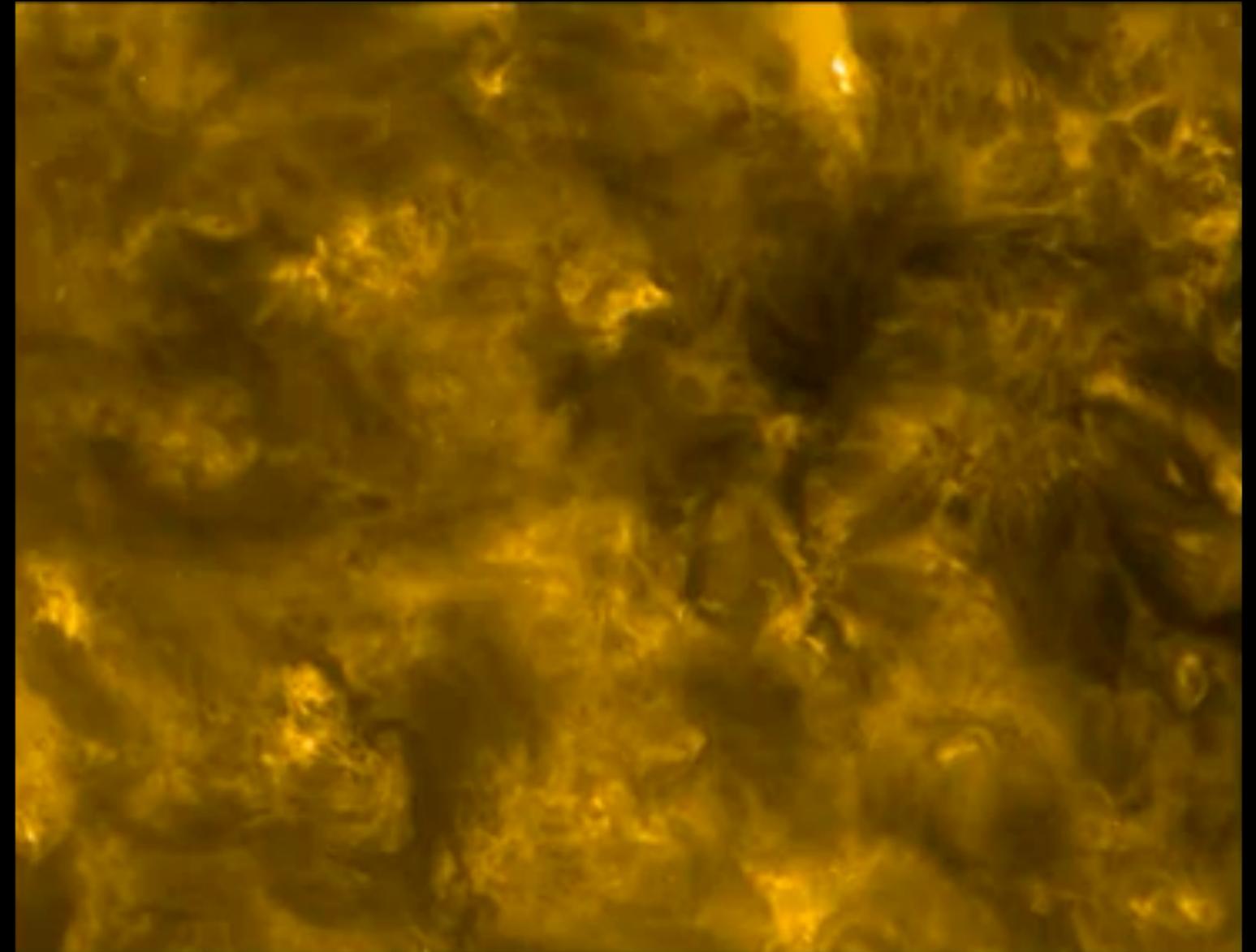
Brightenings: Navdeep K. Panesar+ 2021 ApJL 921 L20

Unresolved impulsive events powering the solar corona.

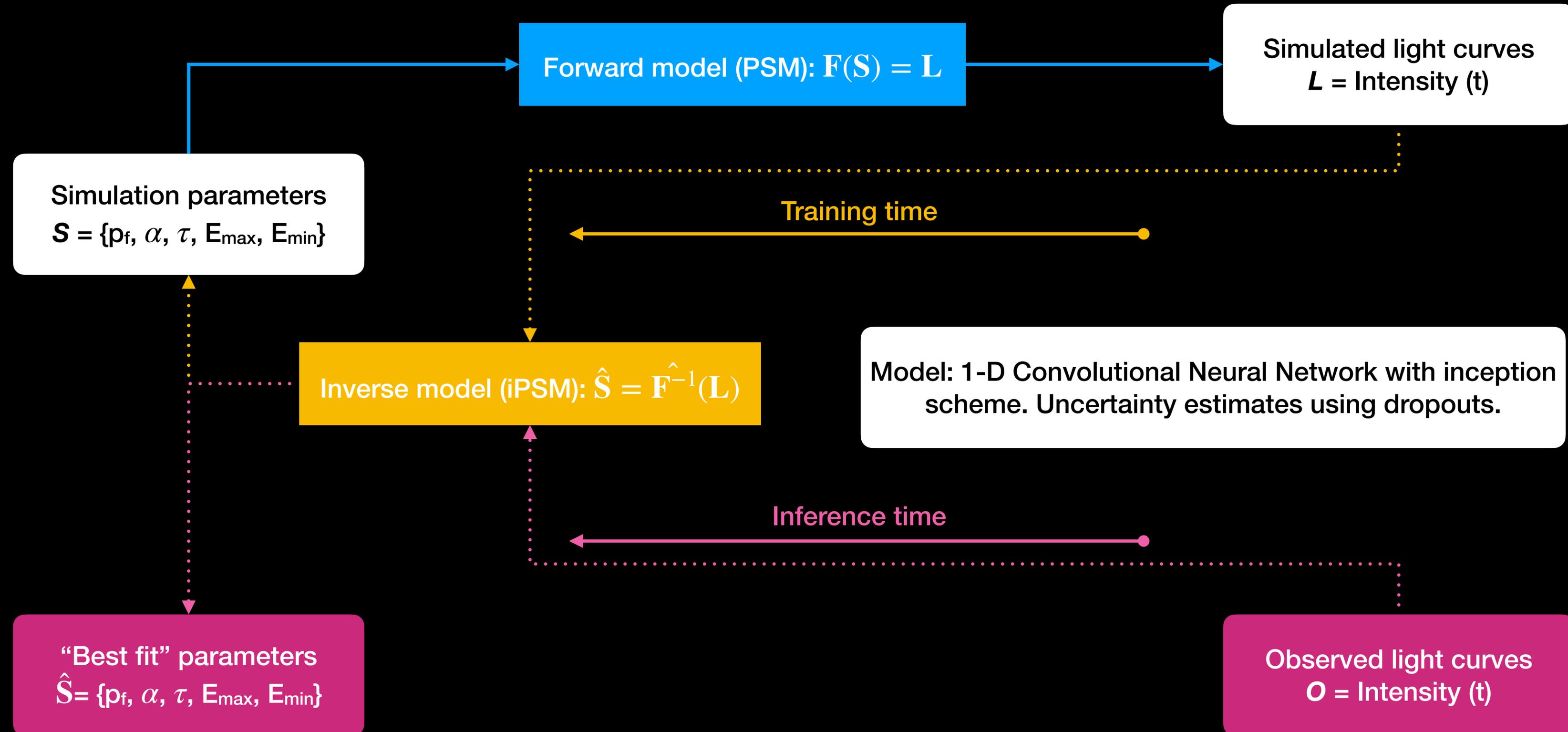
Quiet Sun



- Unresolved impulsive events make up this light curve.
- Statistical constraints on properties of these events => Use statistical simulations to constrain model parameters given data!

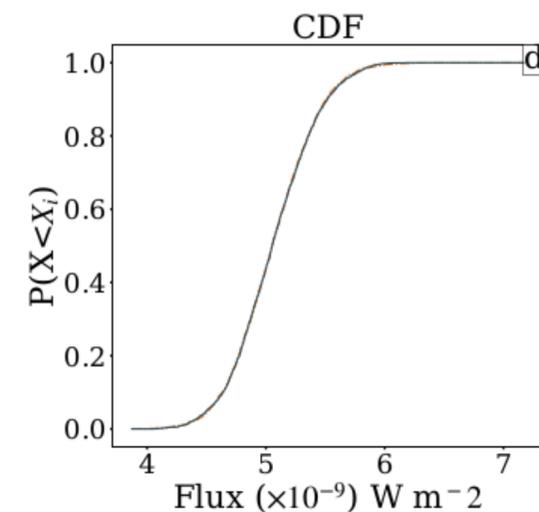
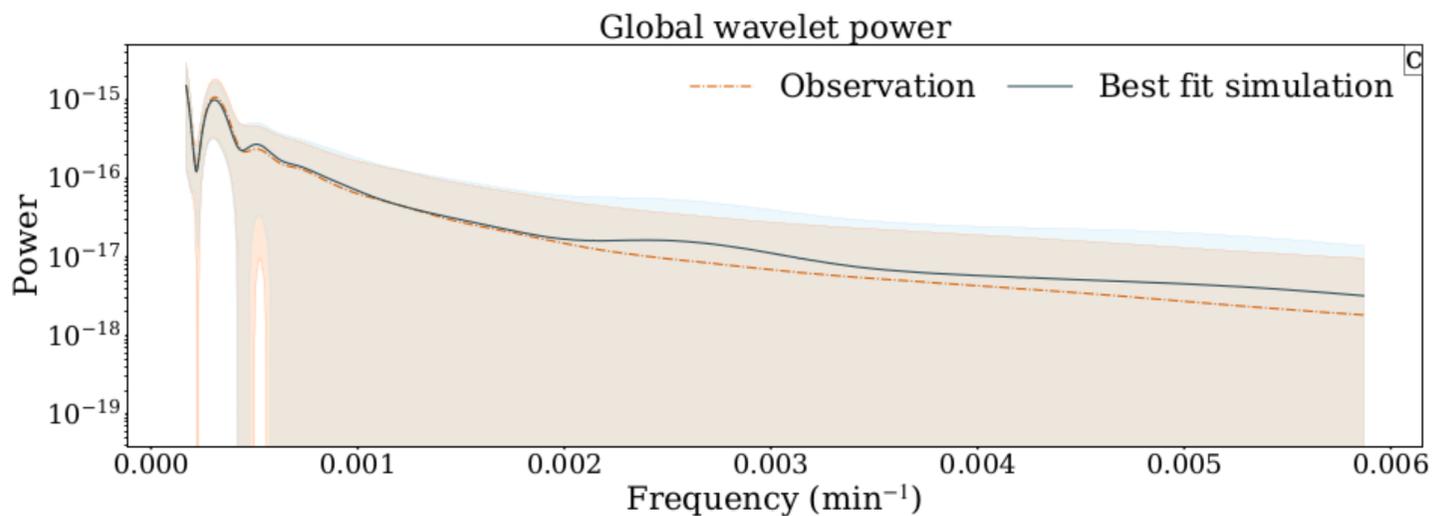
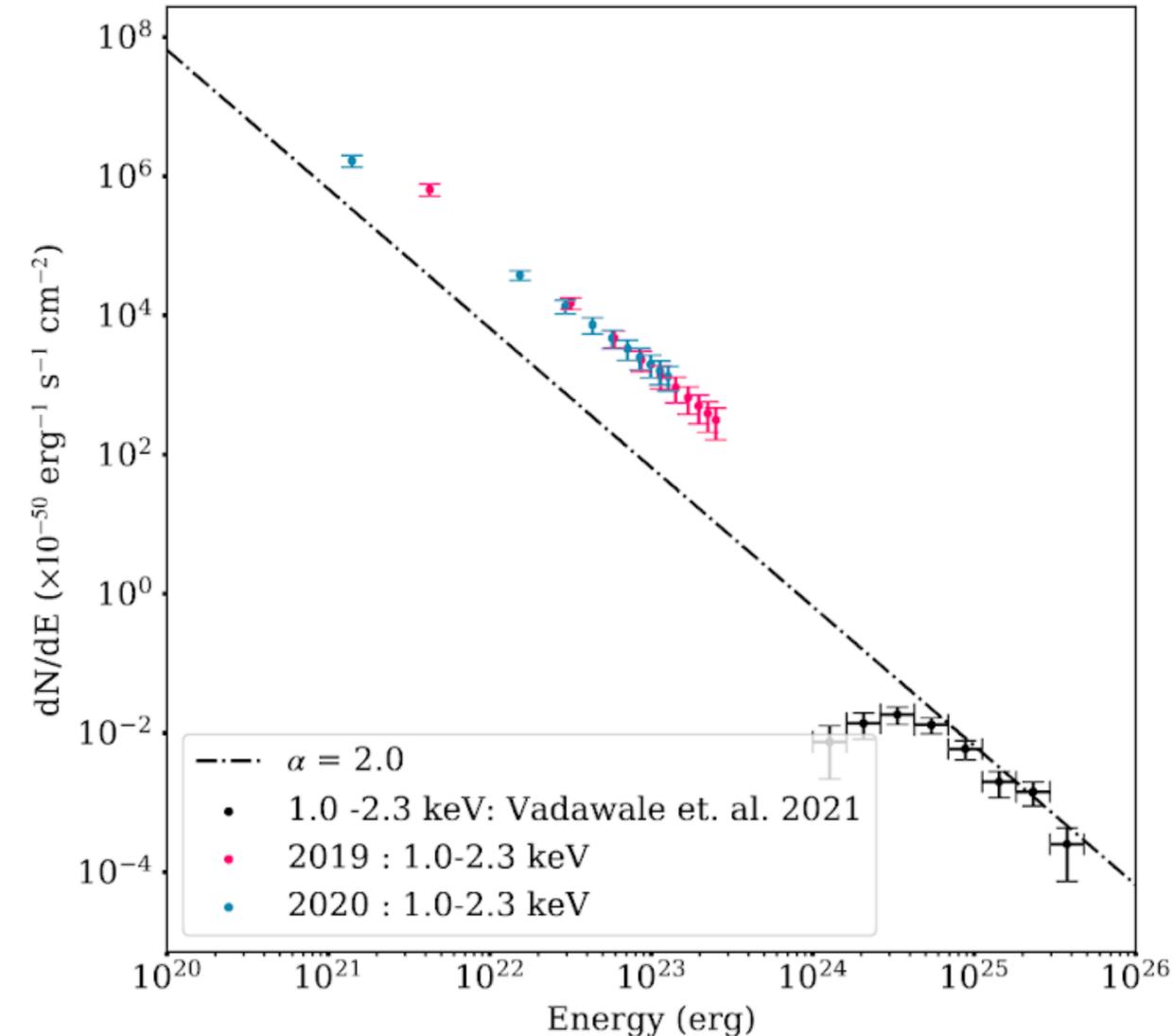
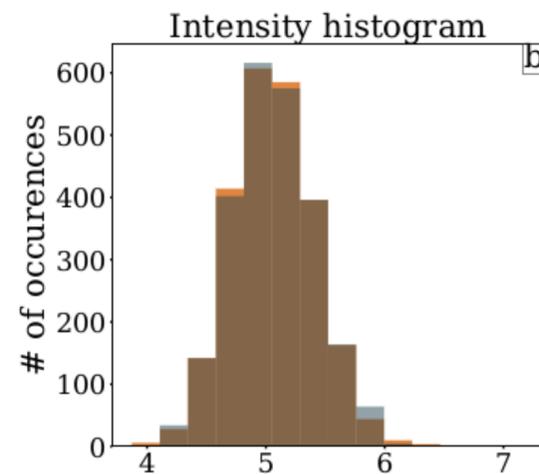
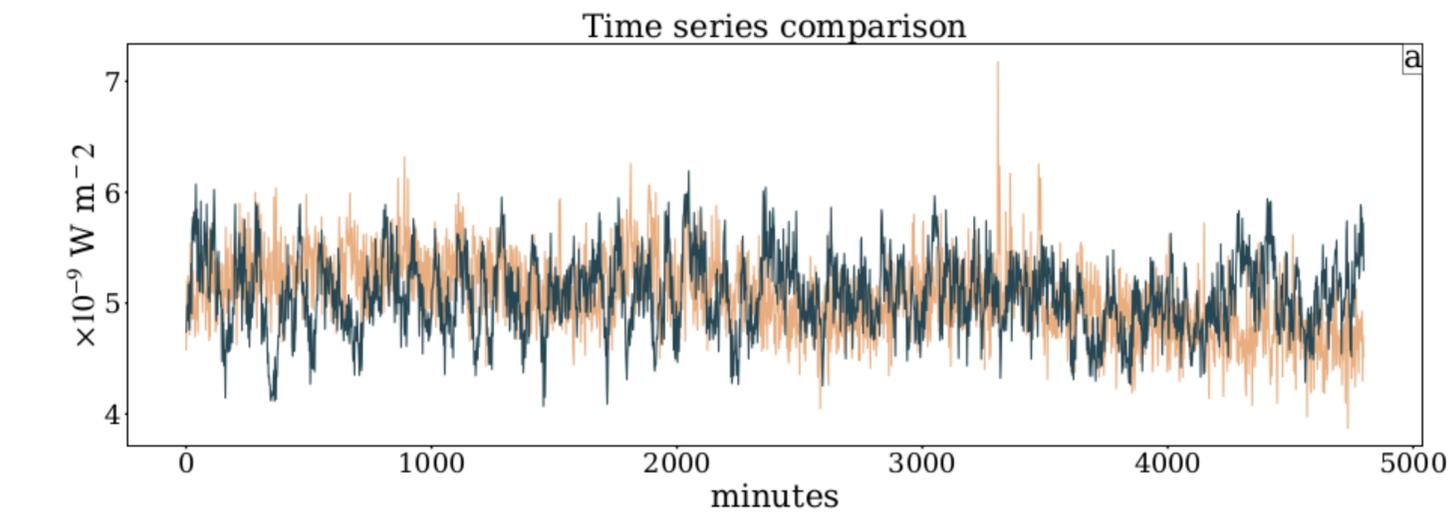


ML inversion scheme for simulation



Powering the solar atmosphere

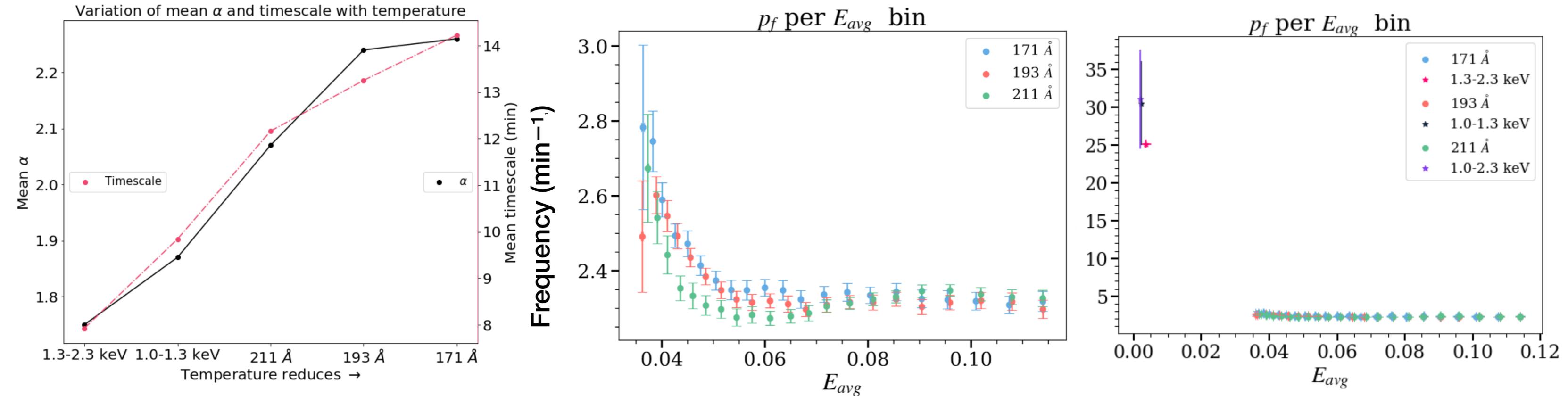
XSM: 2019-10-17 07:43:00 -- 2019-10-21 05:13:00 @ 1.0-2.3 keV



$$\alpha = 1.87 \pm 0.15, \quad y_{\min} = 2.12 \text{e-}12 \text{ W m}^{-2}, \quad p_f = 28.00 \pm 2.16 \text{ min}^{-1}, \quad y_{\max} = 1.26 \text{e-}10 \text{ W m}^{-2}, \quad \tau = 11.80 \pm 1.05 \text{ min},$$

Estimate the parameters of impulsive events for light curves across Extreme Ultraviolet and X-ray energy bands. X-ray events have $\approx 10^{21}$ erg of energy per event — regime of “nano / pico” flares powering the solar corona.

.. we uncover more interesting properties of these events!



1. The more frequent events are, smaller is their amplitude -> presence of energy reservoir.
2. Events get more transient with increasing temperature of plasma.
3. Possible change in energizing mechanism between low and high temperatures?

These results are not possible without ML-based inversion scheme!

Length: $\sim 10^{11}$ m, Time: \sim days

Explainable AI for model evaluation and scientific discovery

Application: Understanding source regions of the solar wind.

Vishal Upendran, Mark Cheung, Shravan Hanasoge, Ganapathy Krishnamurthi 2020. Space Weather, 18, e2020SW002478



Morphological structures in the corona show association with solar wind structures.

Coronal Hole



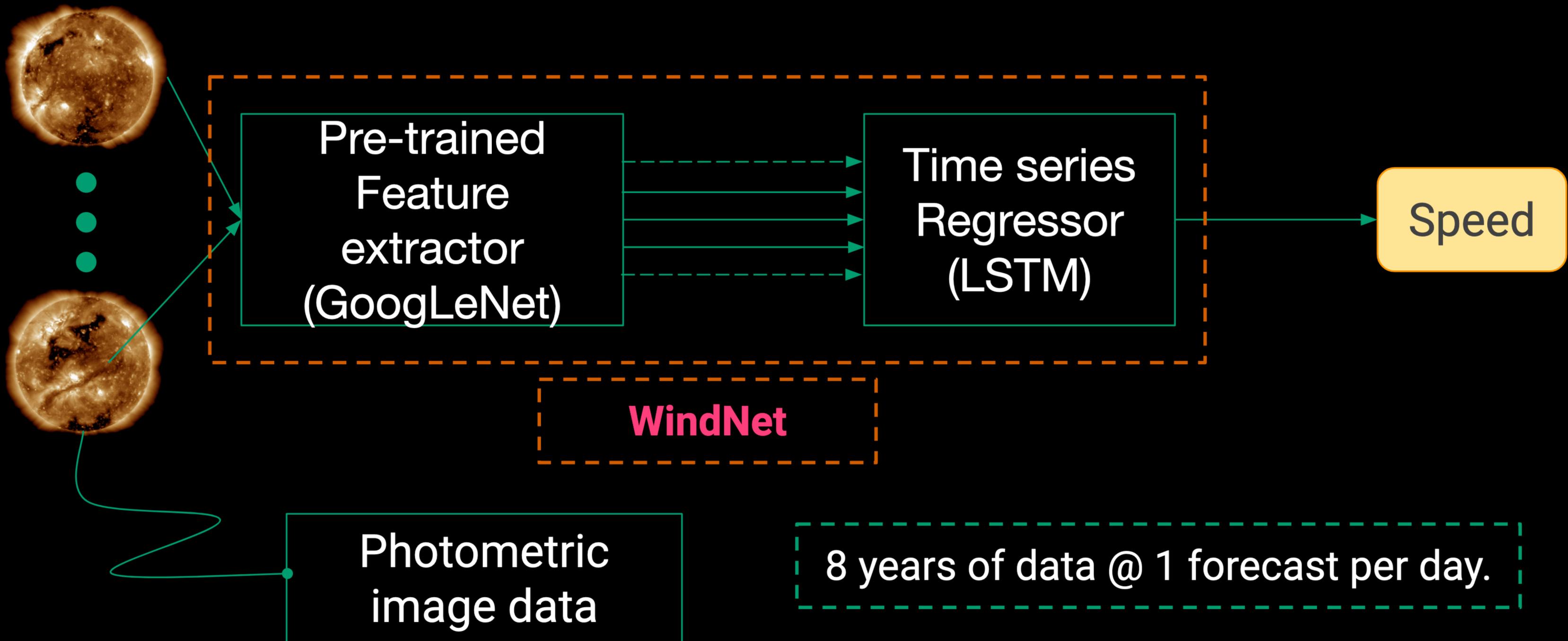
Active Region



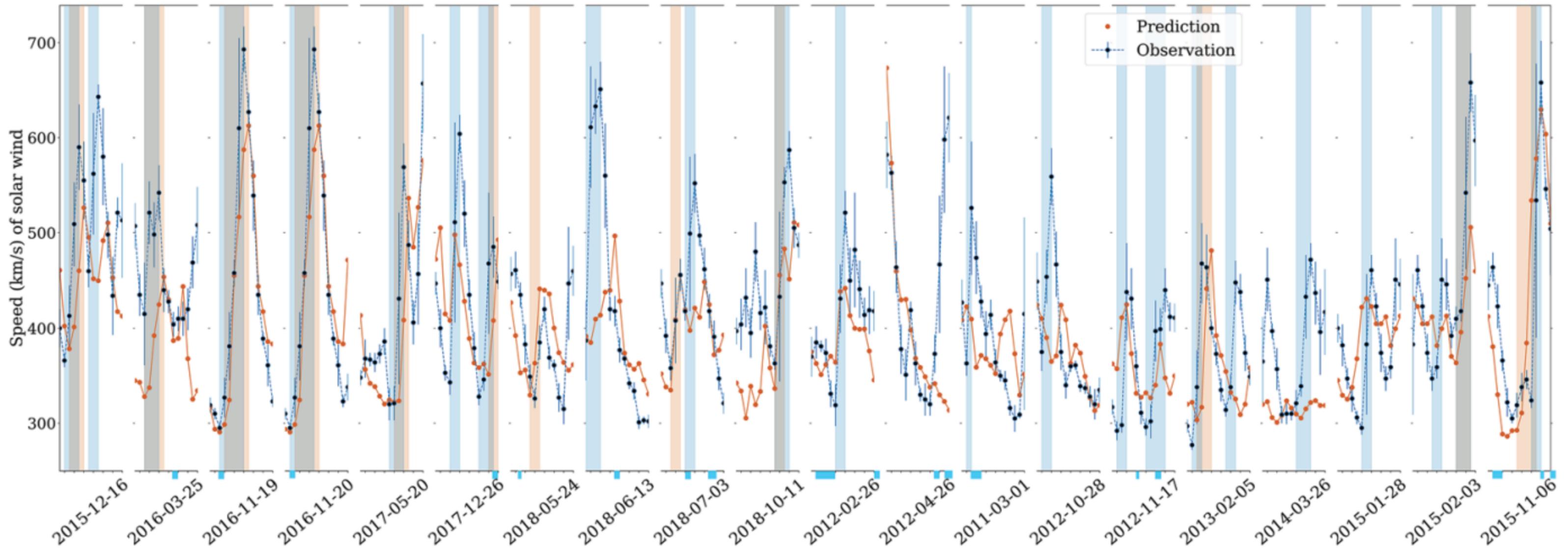
Fast solar wind shows association

Slow solar wind shows association

Translate a sequence of solar images into solar wind measurements



Forecasting solar wind 3-days before it reaches Earth!



Pearson r: **0.61**, RMSE: **76.4 km/s**; Uncertainty normalized squared error: **19.35**

Grad-CAM: Dissecting model to infer solar wind sources!

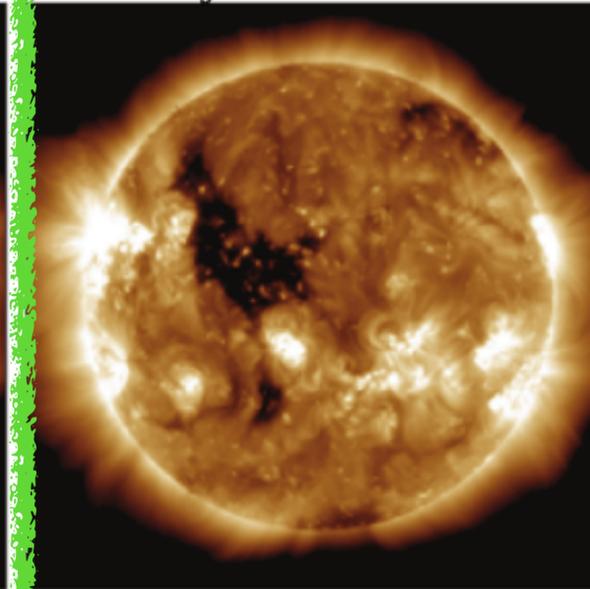
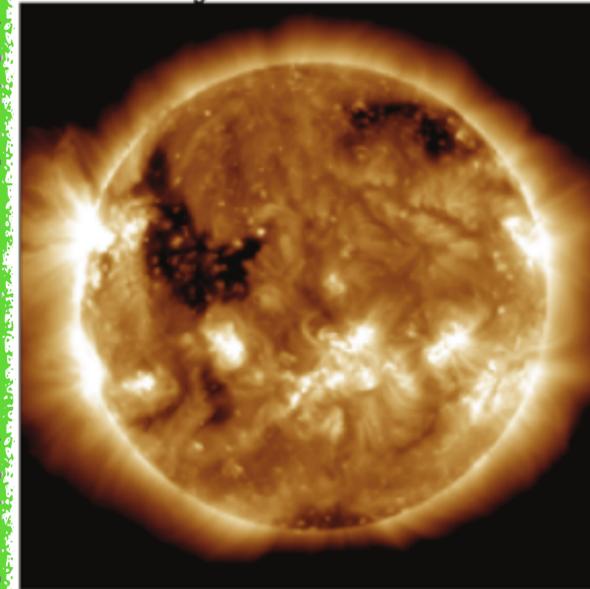
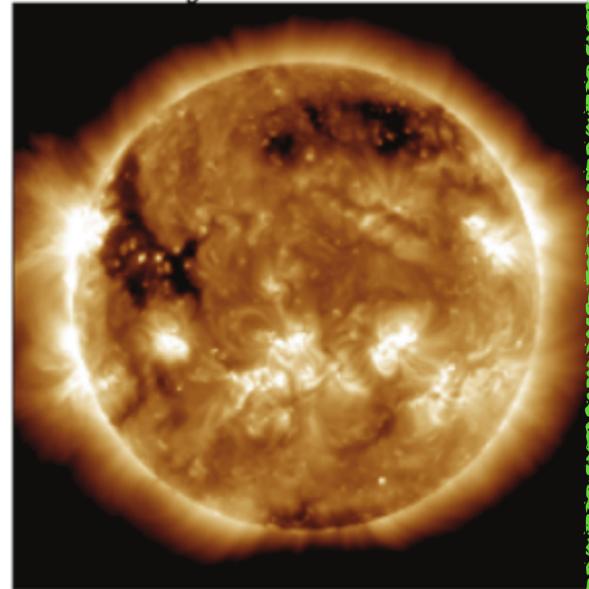
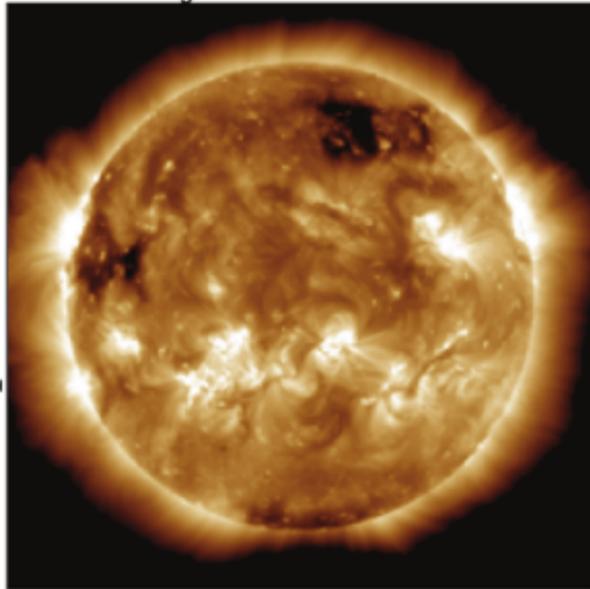
Prediction done for: 2012-06-05

Day: 2012-05-30

Day: 2012-05-31

Day: 2012-06-01

Day: 2012-06-02



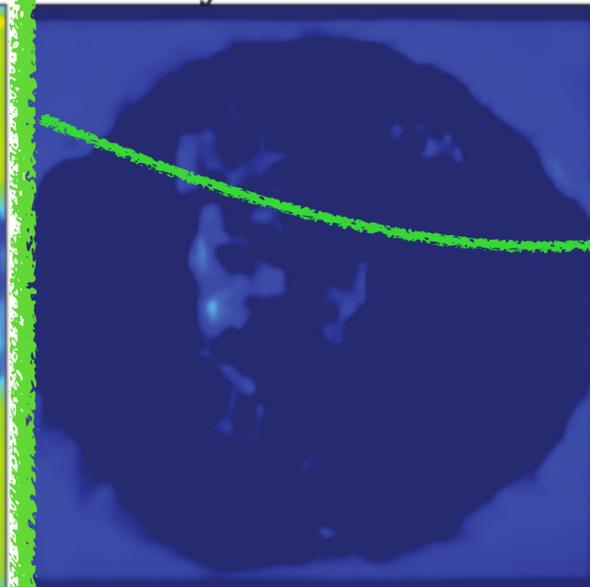
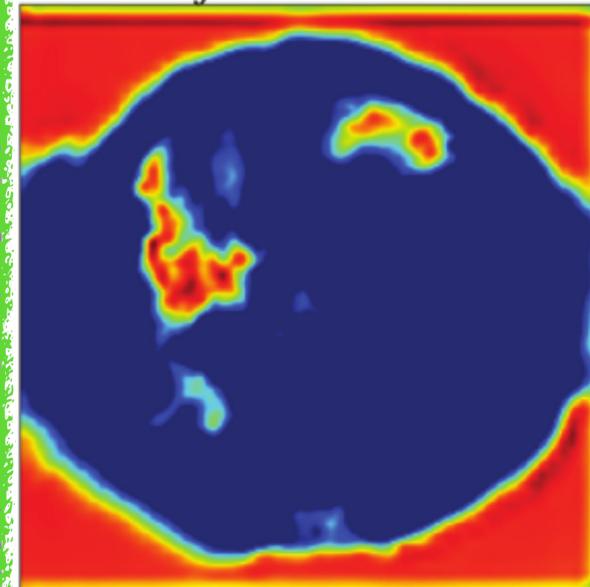
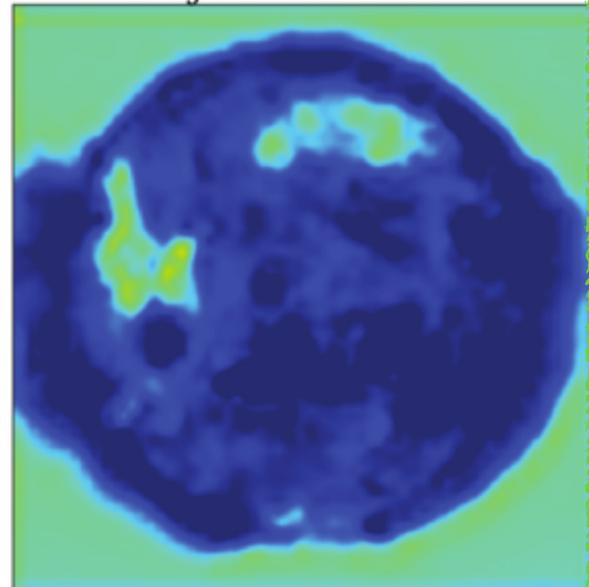
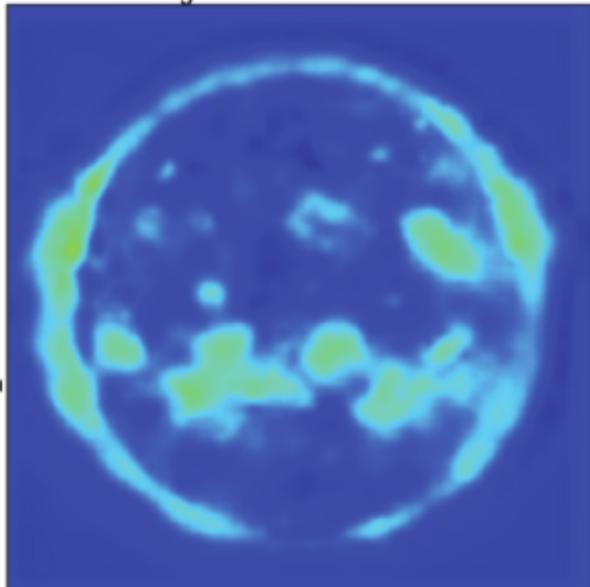
Spatio-temporal coincidence with physics

Day: 2012-05-30

Day: 2012-05-31

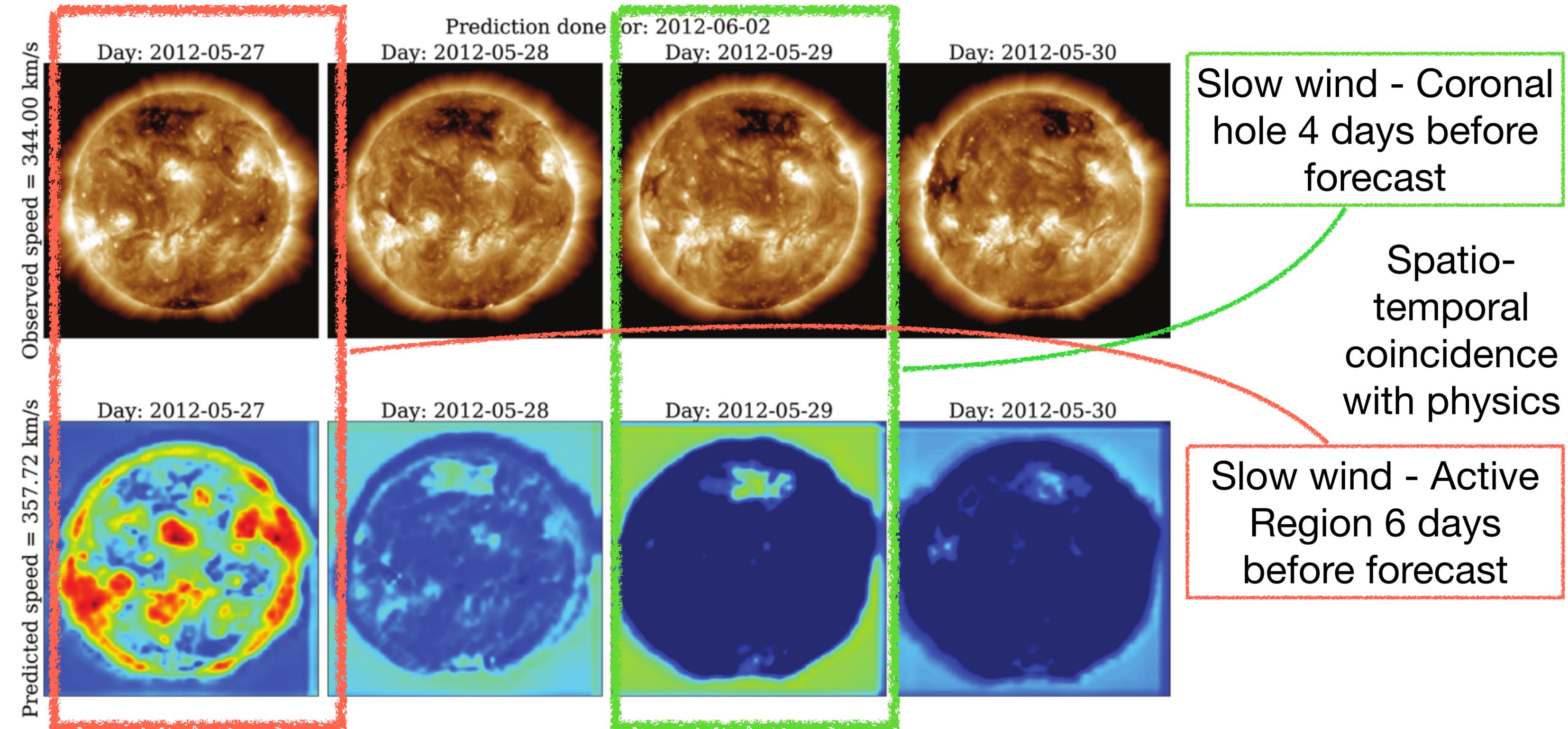
Day: 2012-06-01

Day: 2012-06-02



Fast wind - Coronal Hole association 4 days before forecast

Grad-CAM: Dissecting model to infer solar wind sources!



Statistical

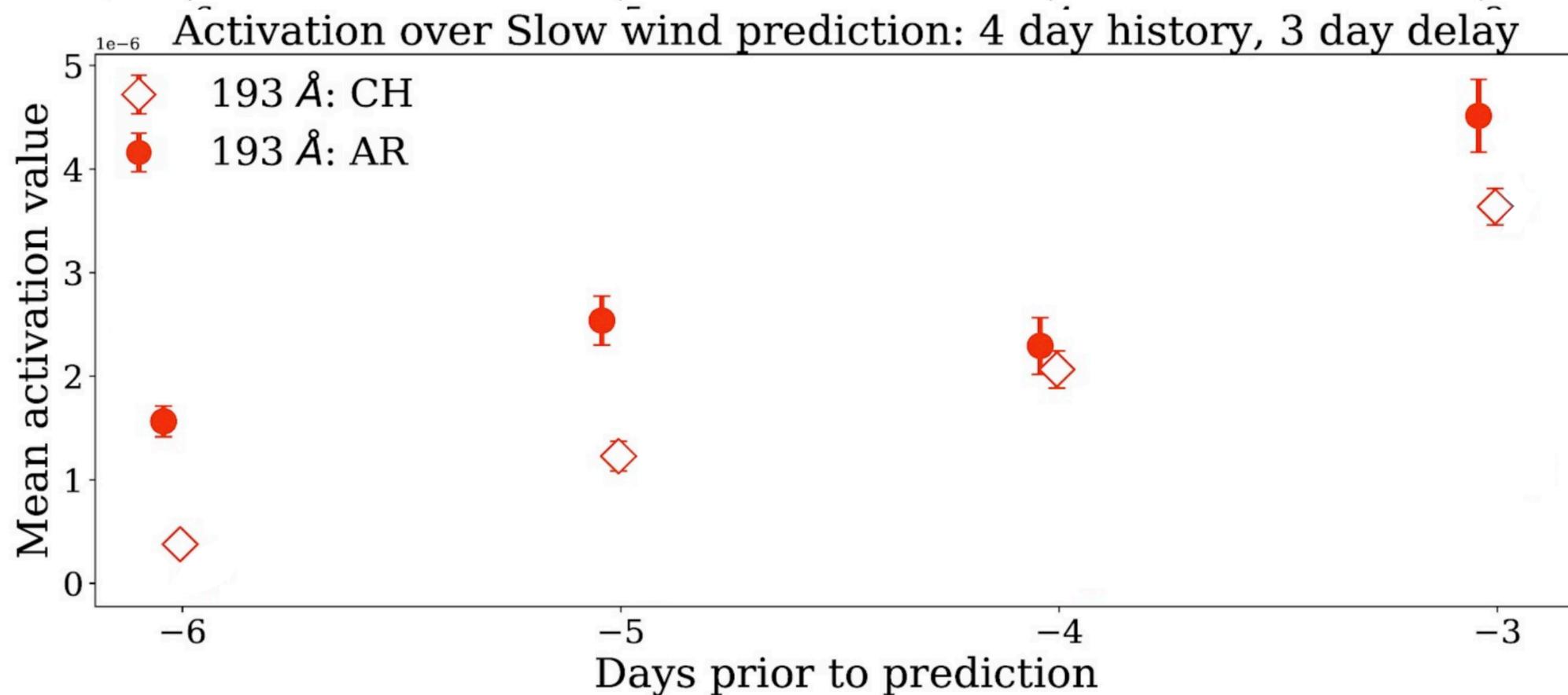
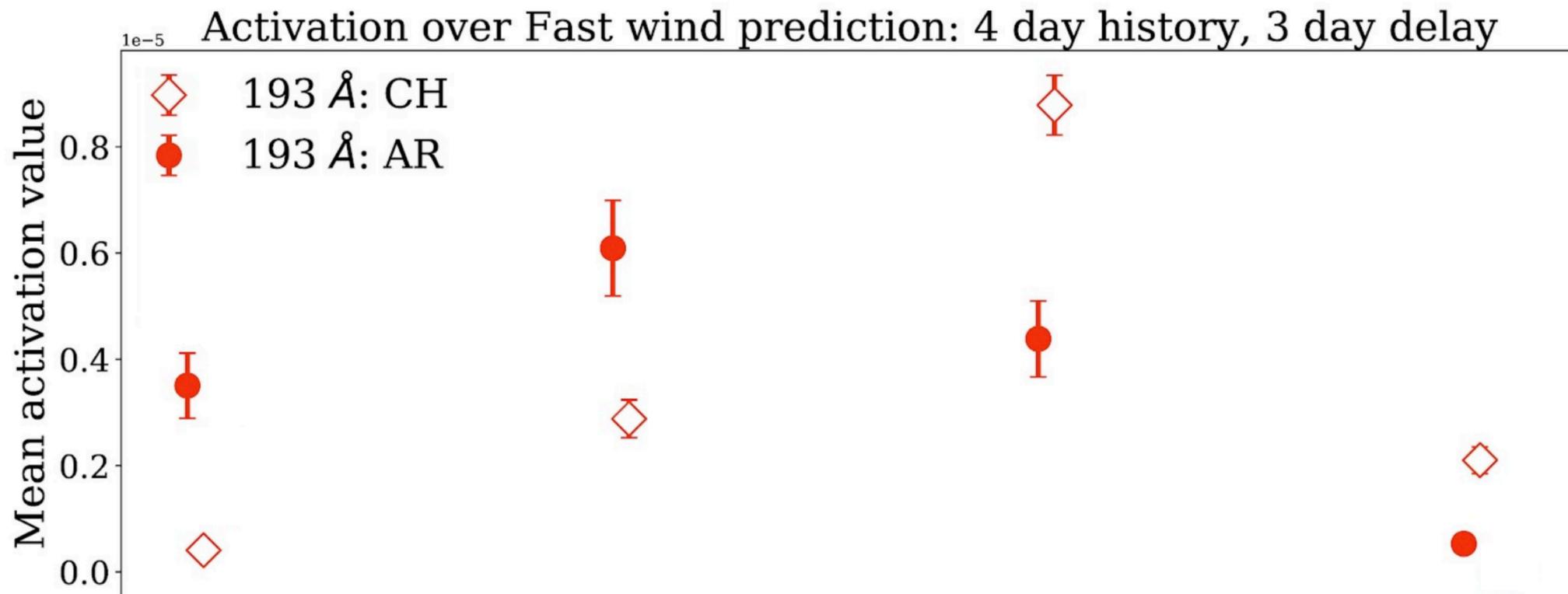
quantification:

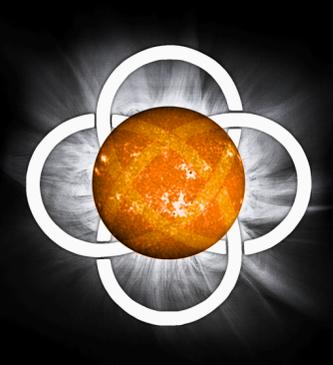
Activation per unit area

1.CH - fast wind consistent with physics.

2.AR - slow wind partially consistent with physics. Time coincidence not seen due to AR lifetime.

3.AR - fast wind needs study.

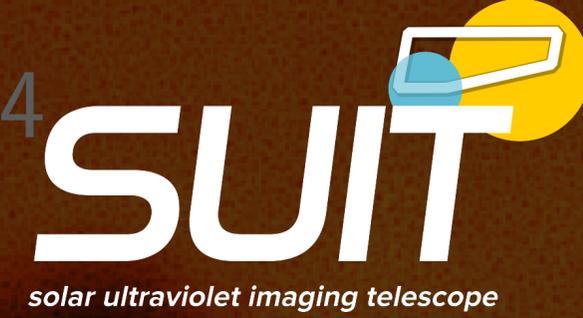




Feb-22 22:29:07

Length: ~ 10⁸ m, Time: ~ mins to hours

NB02 2024-Feb-22 22:29:14



Explainable AI for model evaluation and scientific discovery

Understanding pre-flare signatures in flaring Active Regions

306 2024-Feb-22 23:44:07

NB07 2024-Feb-22 23:54:48

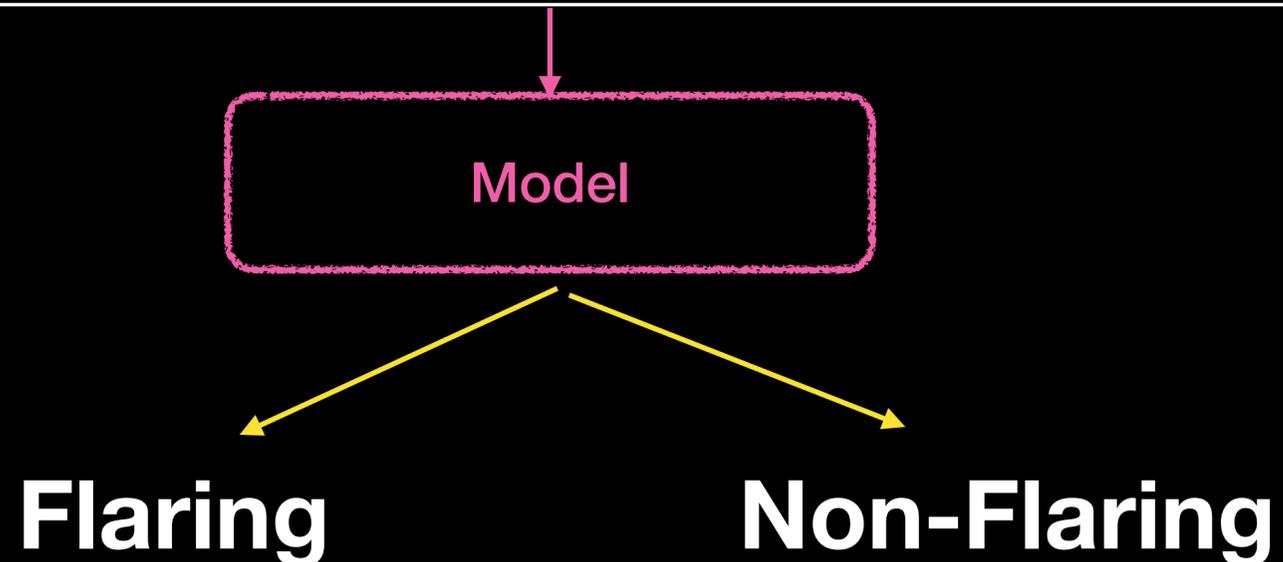
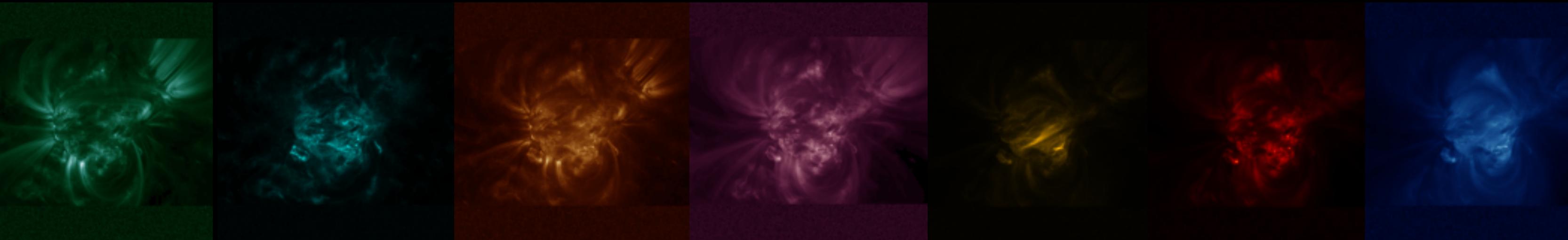
With Linn Abraham, Durgesh Tripathi, *The SUIT Team*
(In-prep)



PhD thesis of Mr. Linn Abraham as part of ISRO - RESPOND project on interpretable solar flare forecasting algorithm development

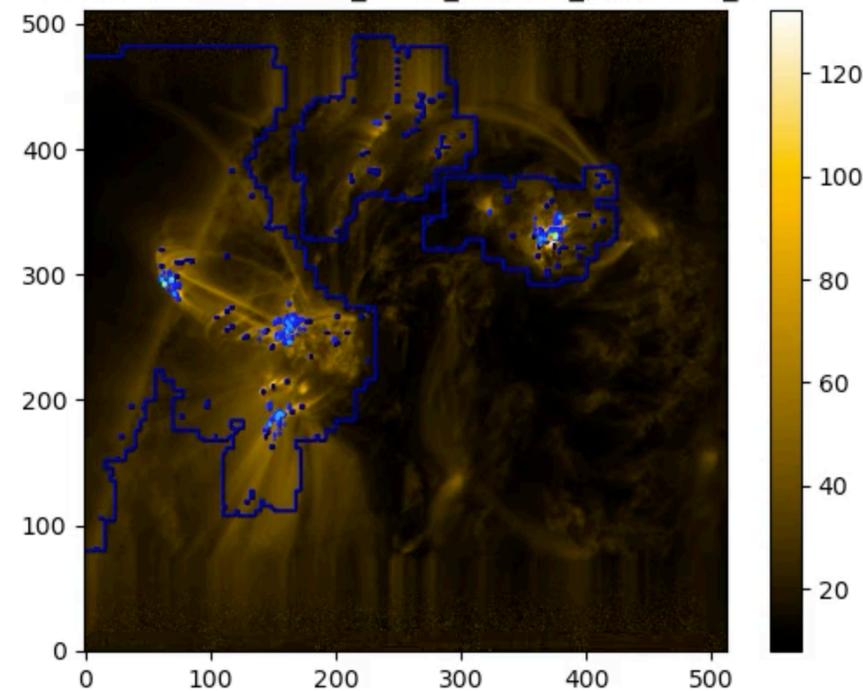
Our approach to investigation:

Can we classify multi-wavelength AIA images of Active Regions in Flaring / Non-flaring regions to understand pre-flare signatures?

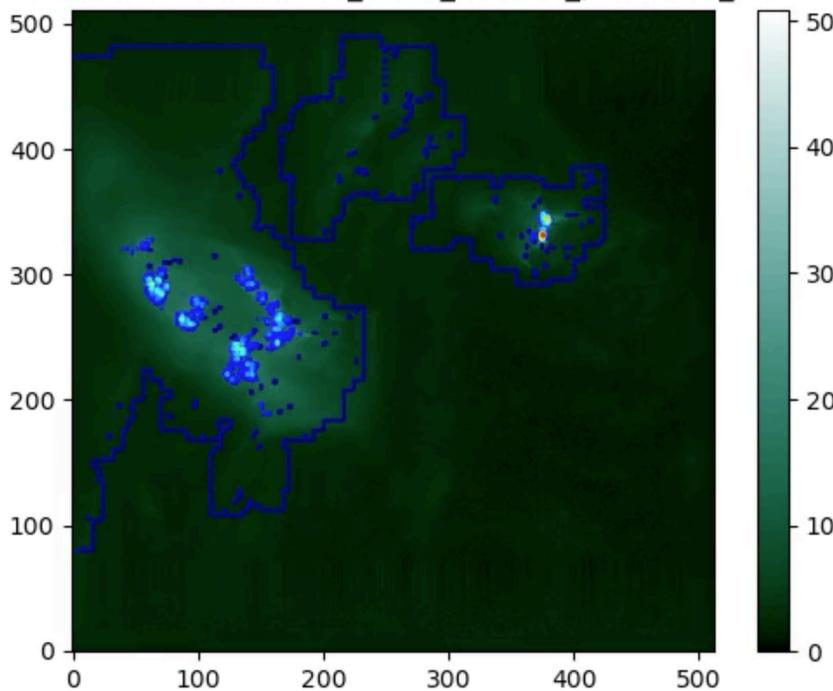


Our work: Classify AR patches to Flaring

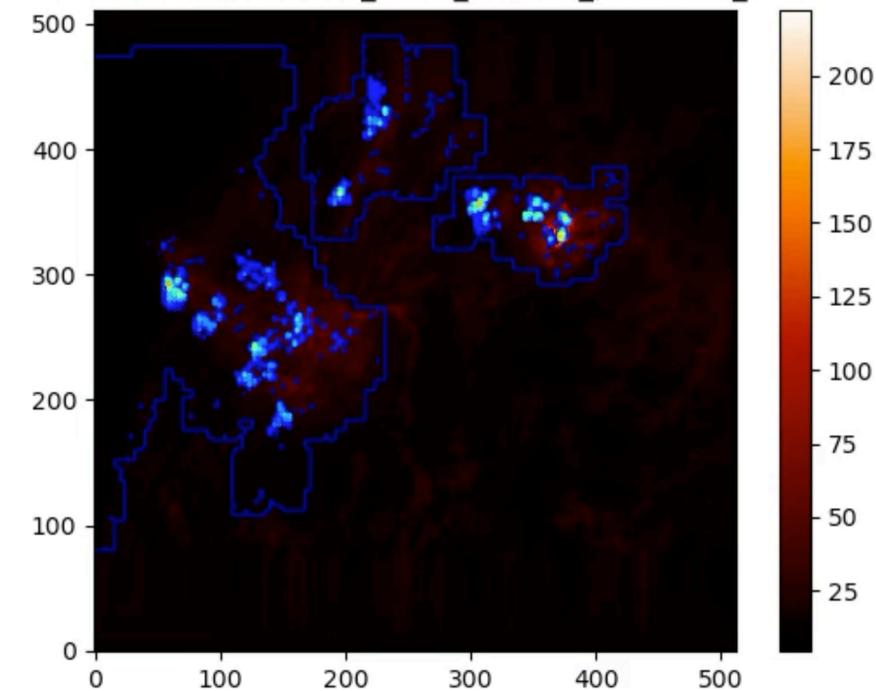
2012-03-04T15:42:02Z_AARP_Id:1449_passband_171



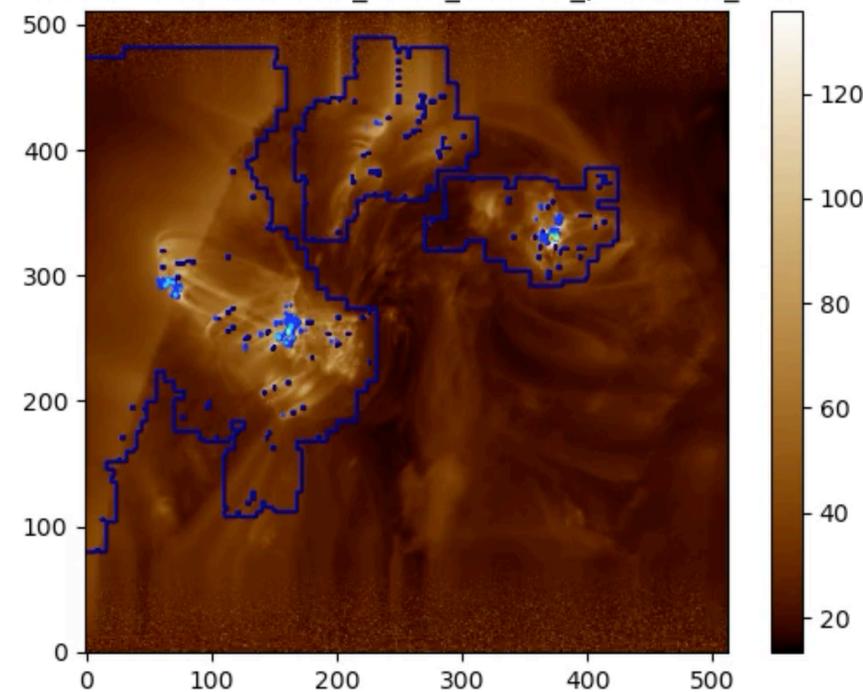
2012-03-04T15:42:02Z_AARP_Id:1449_passband_94



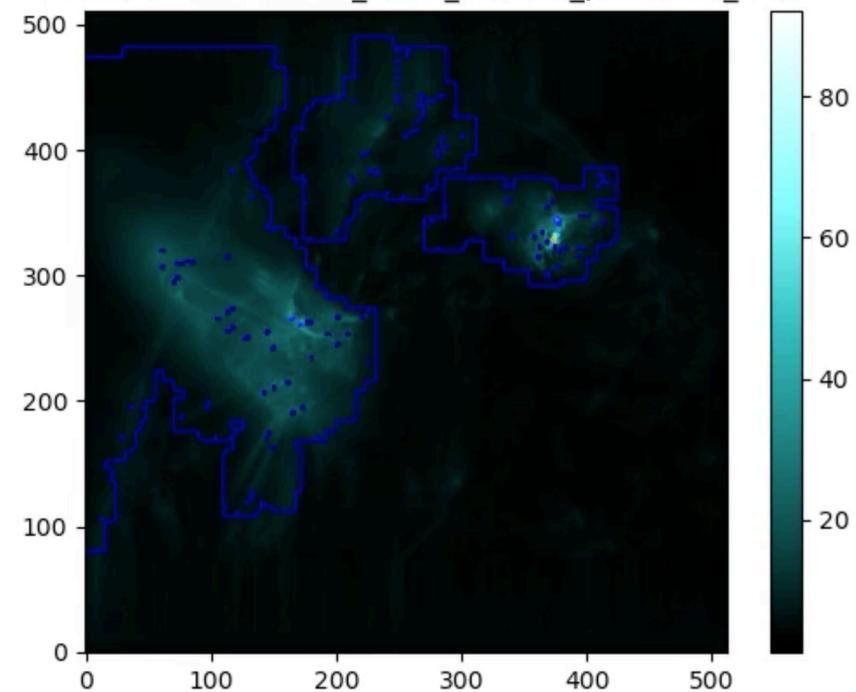
2012-03-04T15:42:02Z_AARP_Id:1449_passband_304



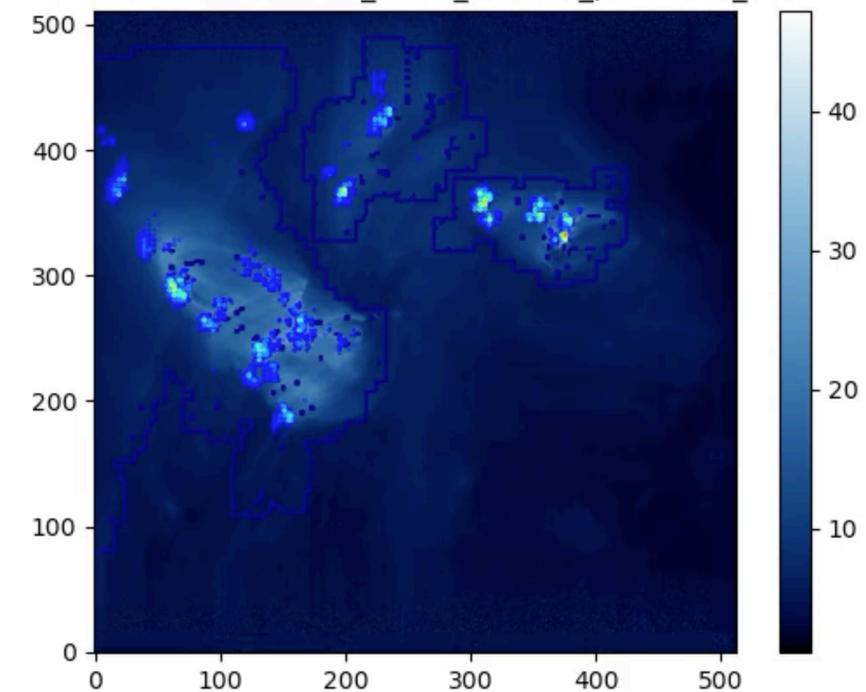
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2012-03-04T15:42:02Z_AARP_Id:1449_passband_131

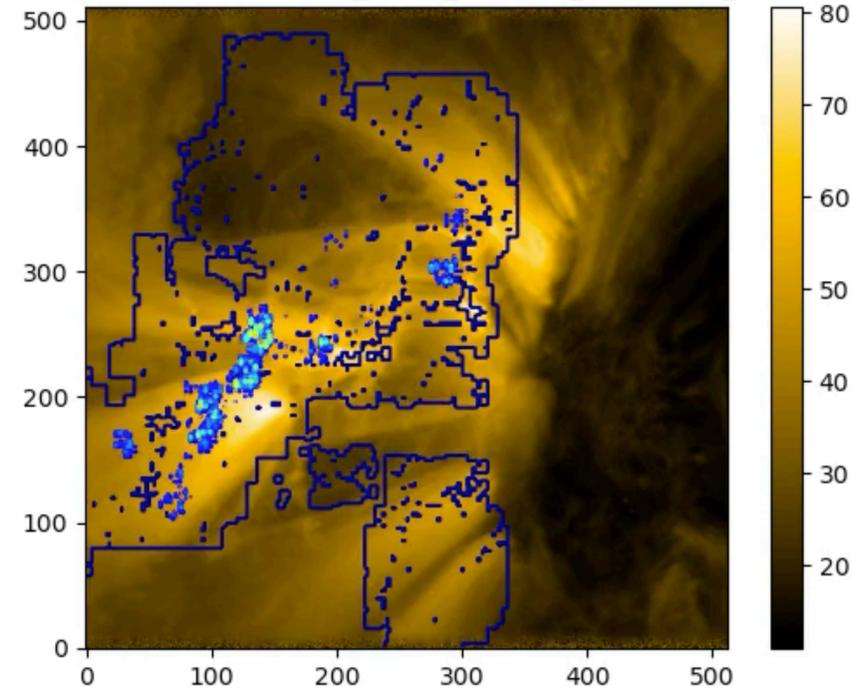


2012-03-04T15:42:02Z_AARP_Id:1449_passband_335

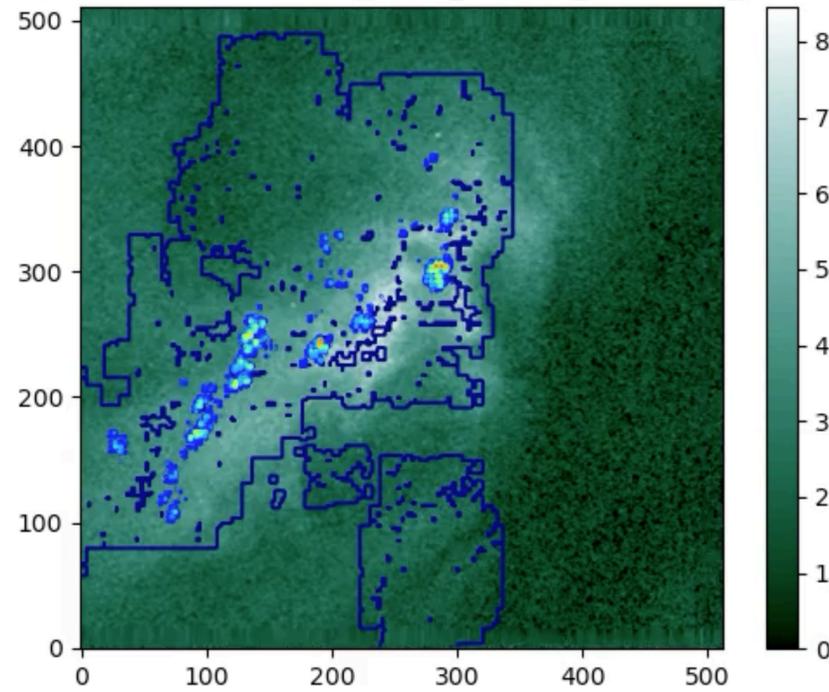


Classify AR patches to non-Flaring

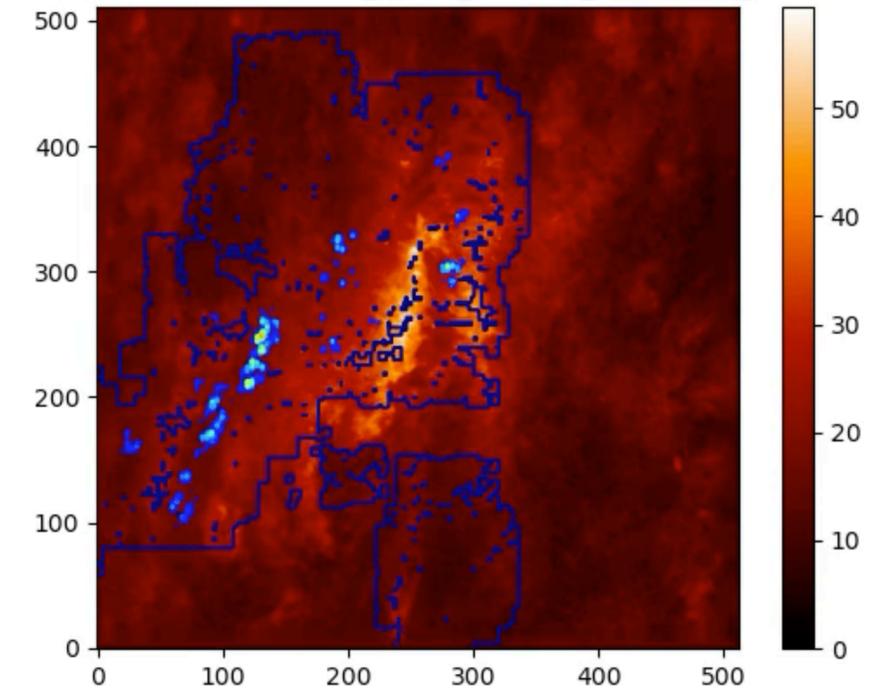
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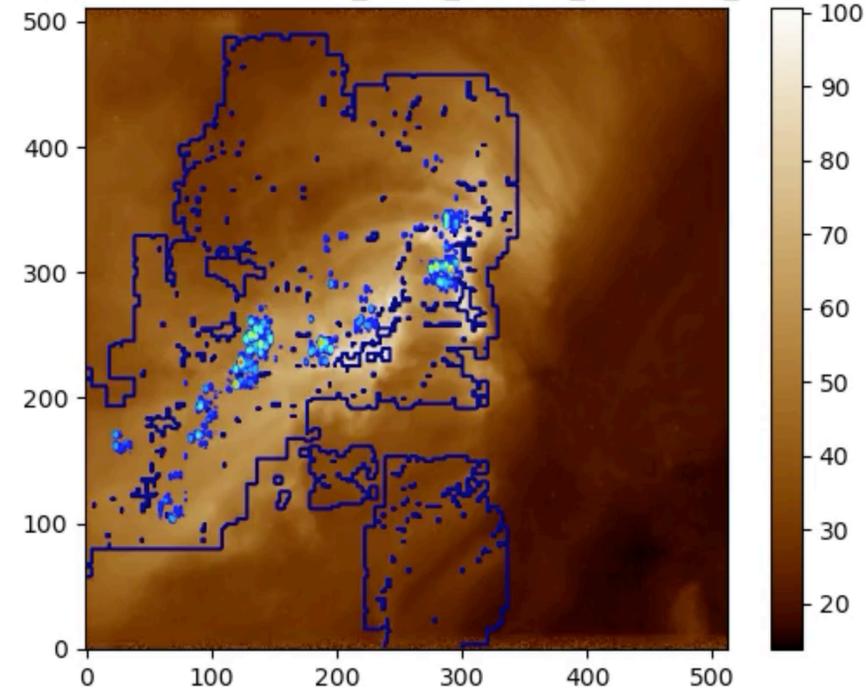
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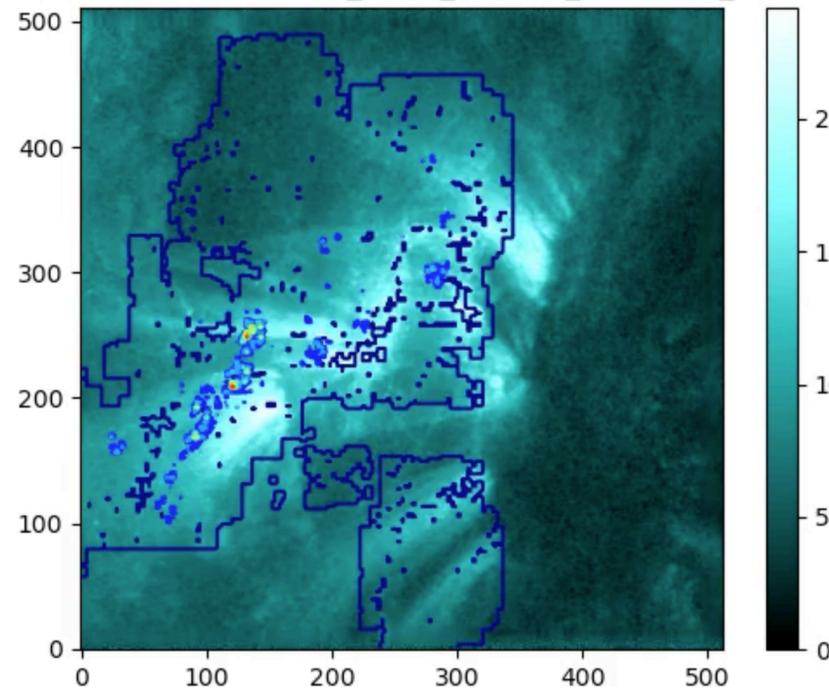
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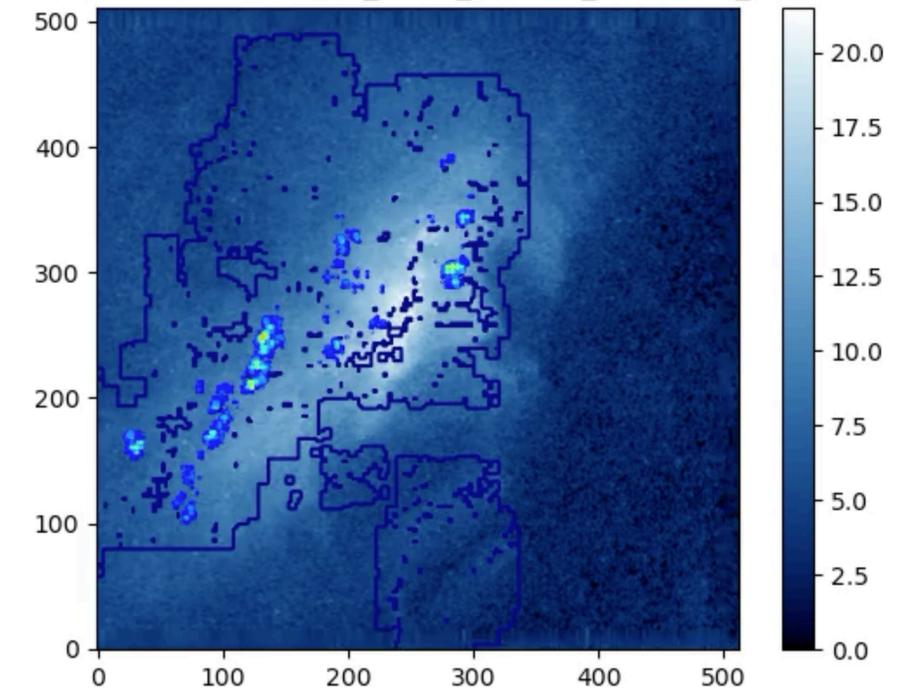
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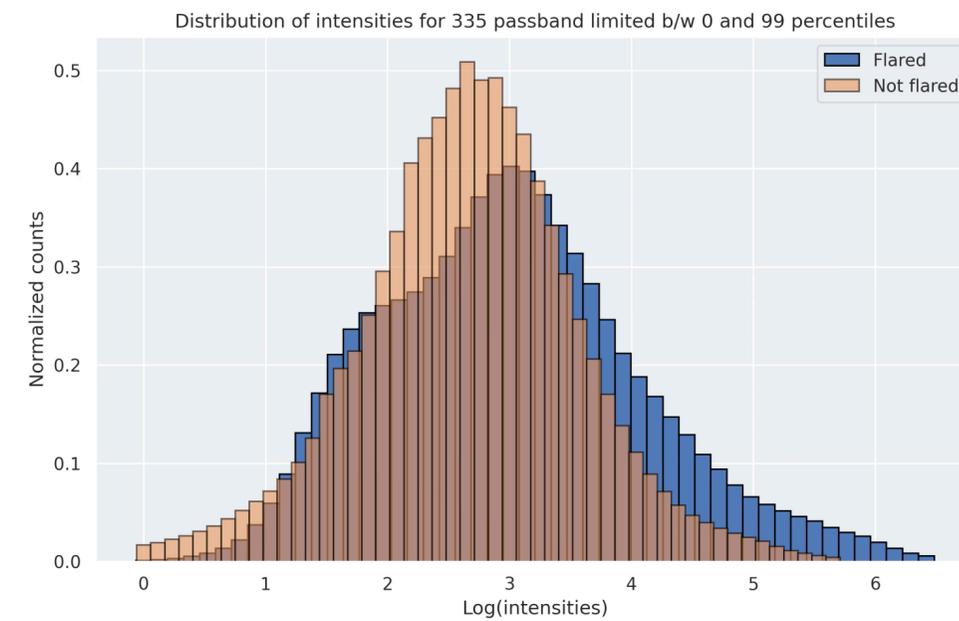
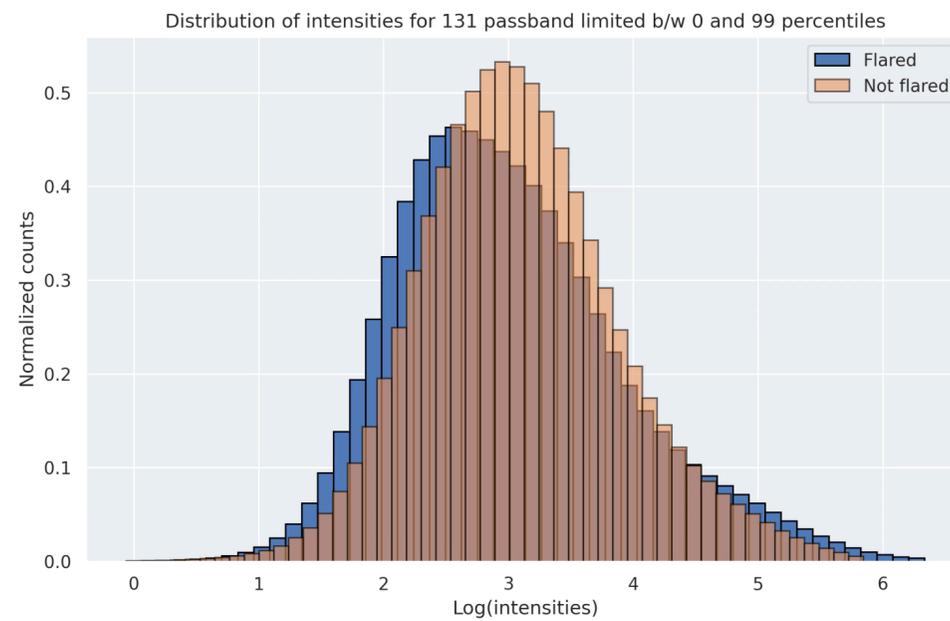
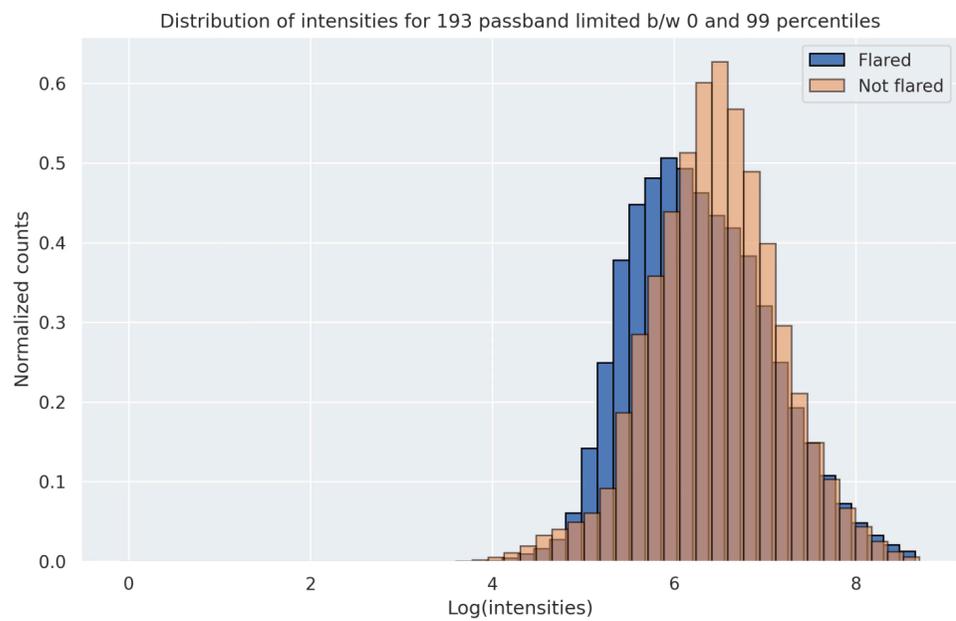
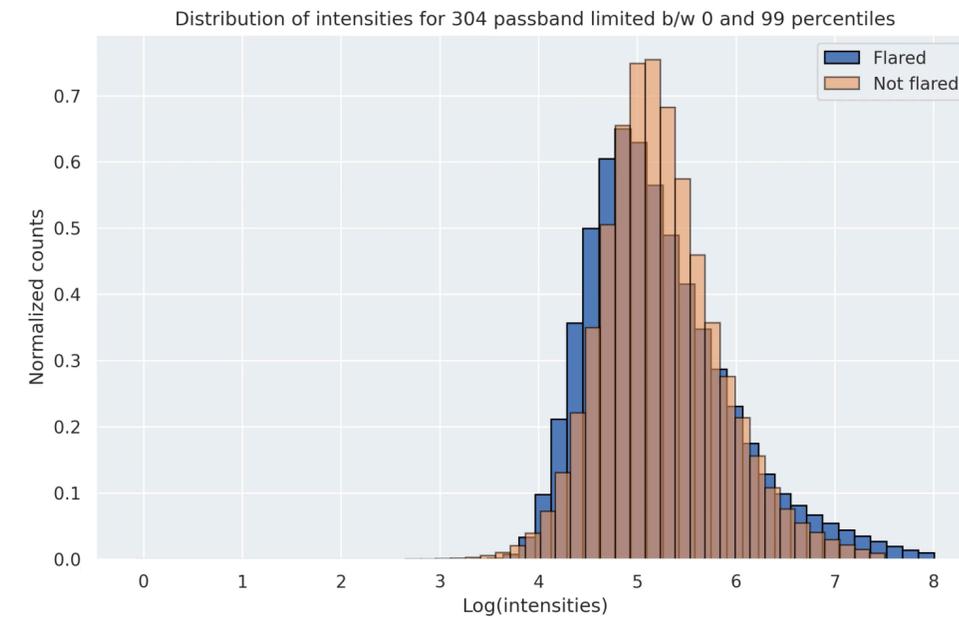
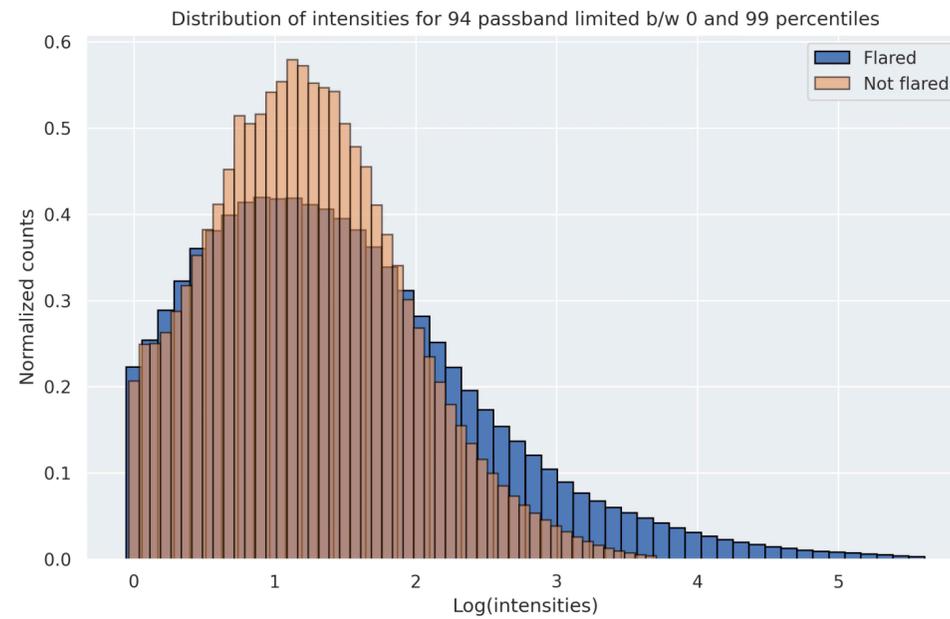
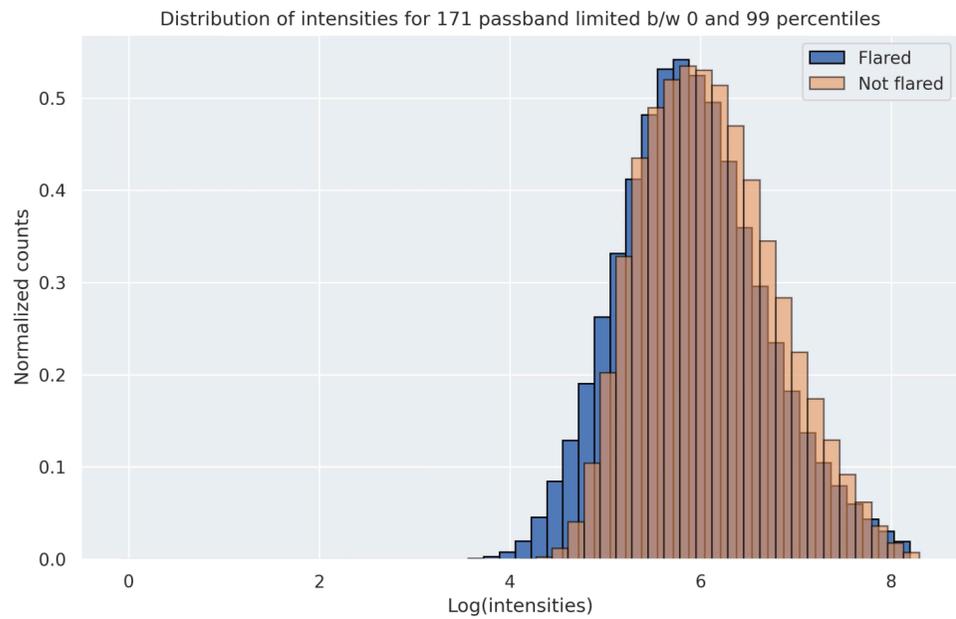


2018-05-09T15:41:59Z_AARP_Id:7256_passband_131



2018-05-09T15:41:59Z_AARP_Id:7256_passband_335





Length: $\sim 10^6$ m, Time: \sim mins to hours to years



The DAGGER Team



Google Cloud

Physics-incorporated deep learning for sparse sampled regression

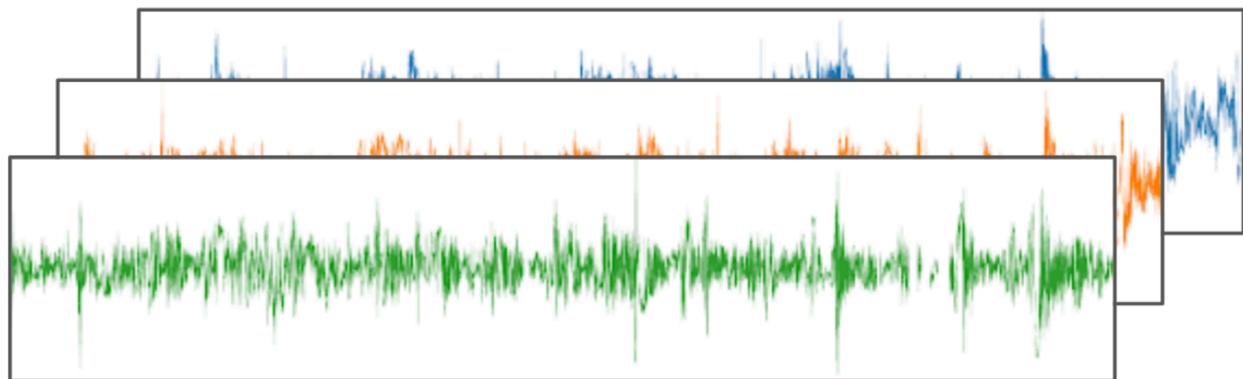
Application: Ground magnetic field perturbation forecasting

Upendran+ 2022. Space Weather, 20, e2022SW003045



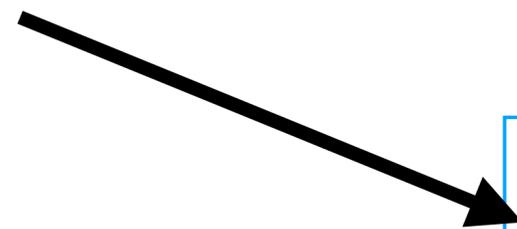
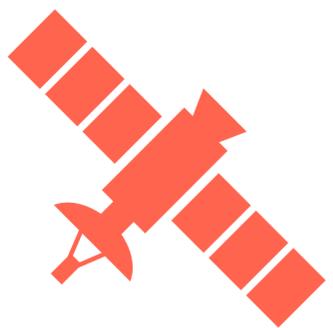


Translating from time series of a single point in space to global effect on the Earth



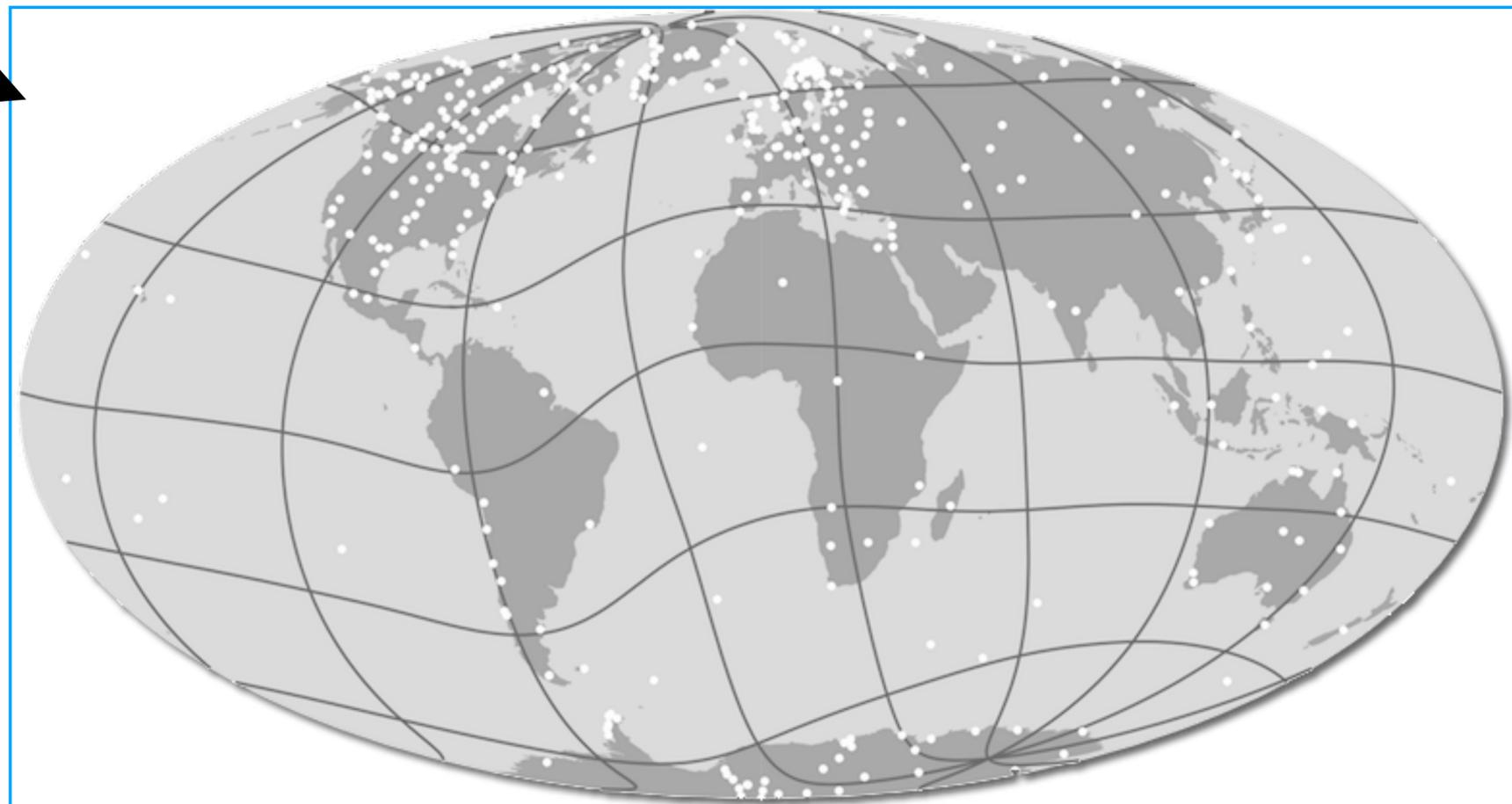
Solar wind measured in space, just outside of Earth's magnetosphere

Solar wind

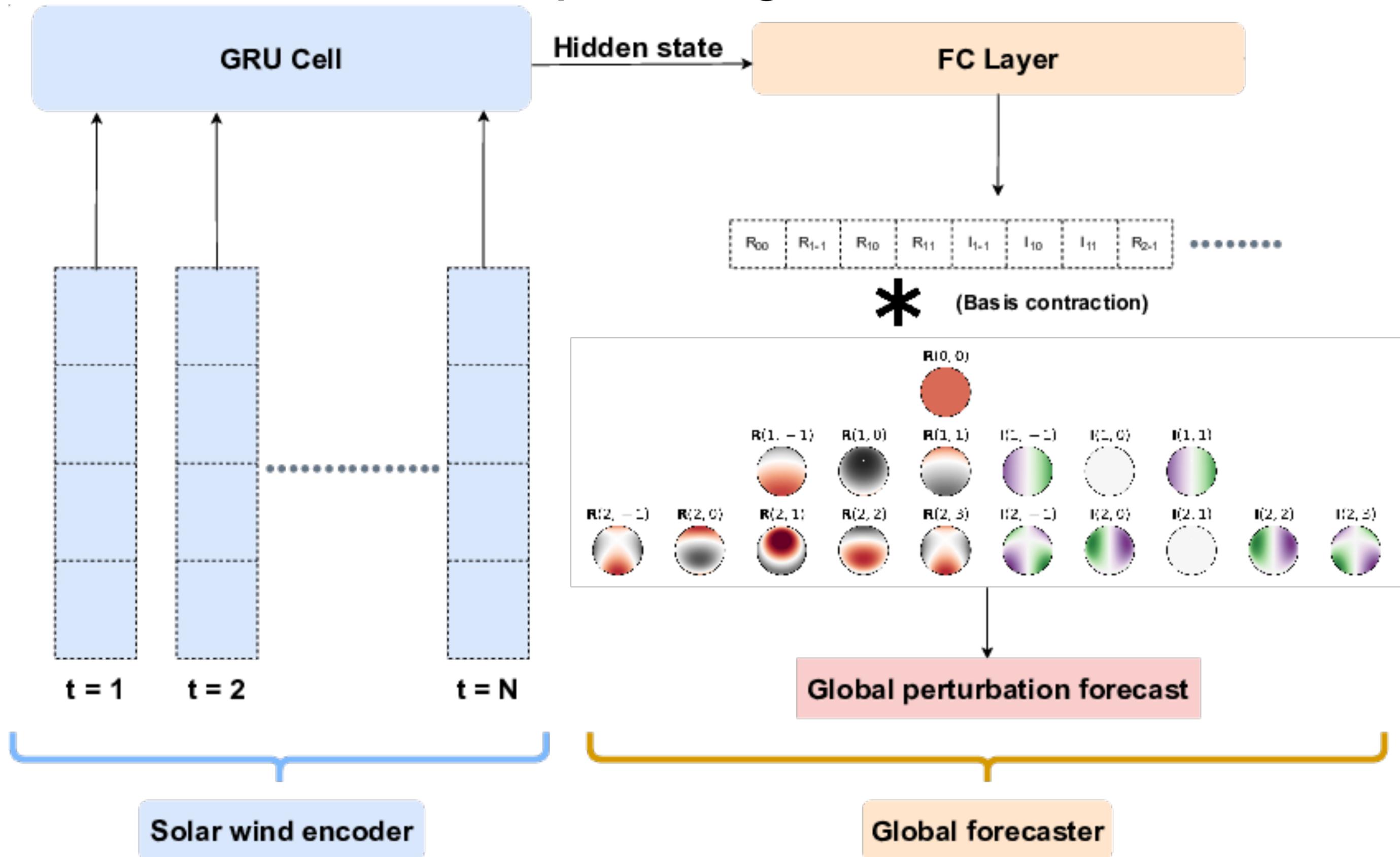


Fluctuations in Earth's magnetic field (δb_i) sampled at various locations on the ground

Stations on the Earth

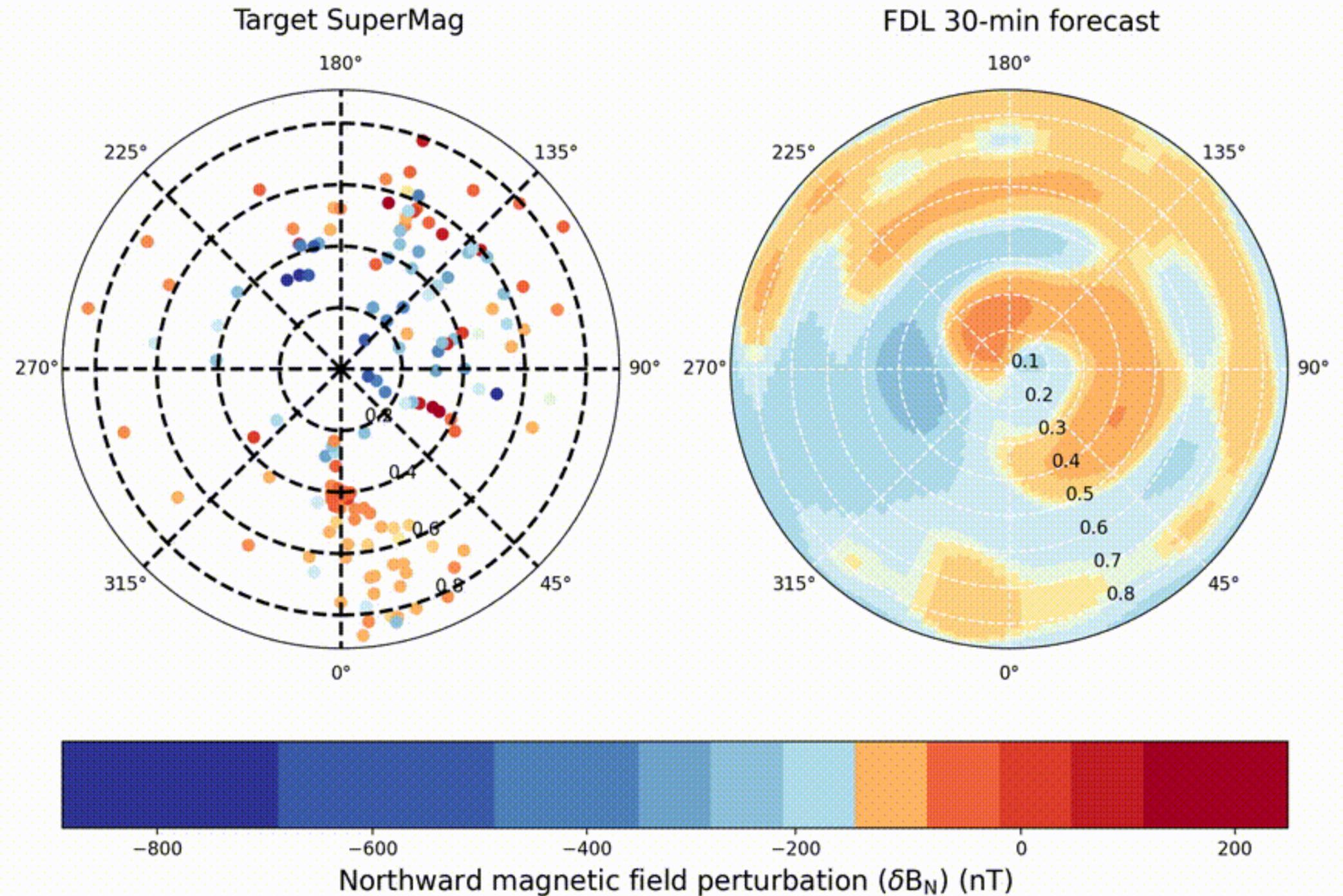


A deep learning solution!



Forecasting across the globe at 1-minute cadence

Capturing global and local scales at high time cadence with the same model : first such model!

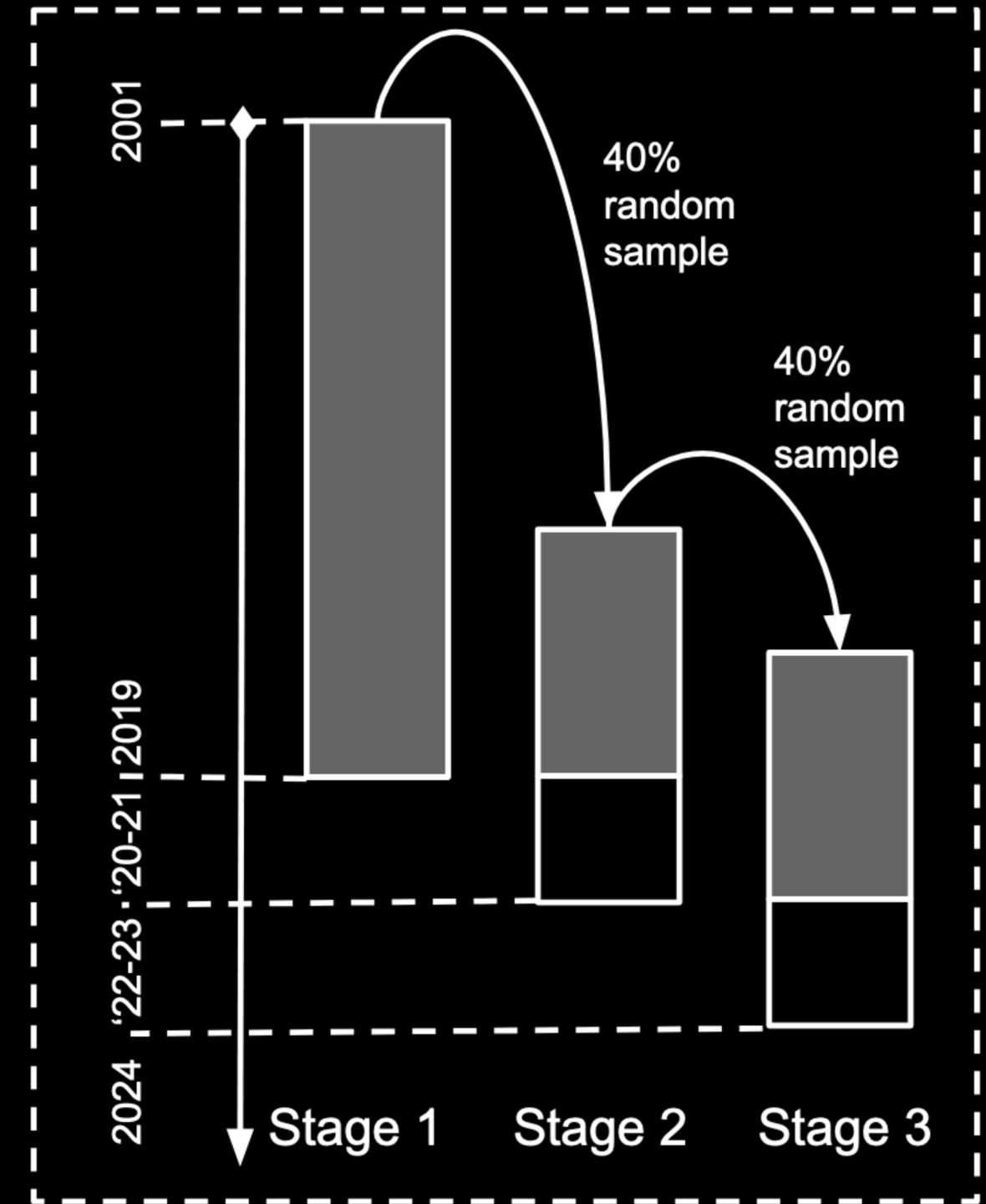


DAGGER CL: Continual Learning

Domain Incremental Learning

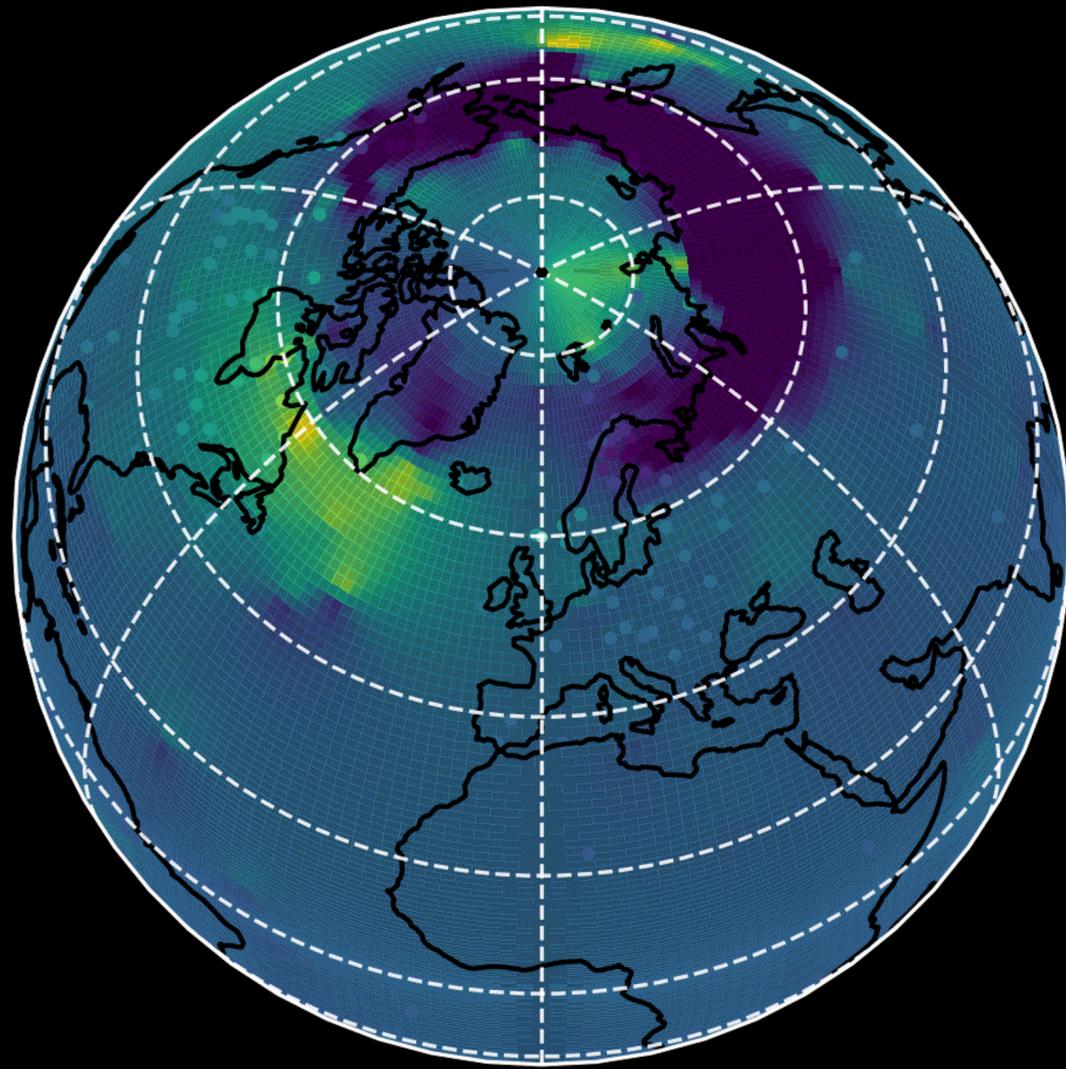
- Data resampling strategy
 - Experiments conducted in 3 stages
 - Data from both ACE and DSCOVR
 - Repeated data for handling catastrophic forgetting
 - Random samples to reduce bias accumulation

In-prep work!

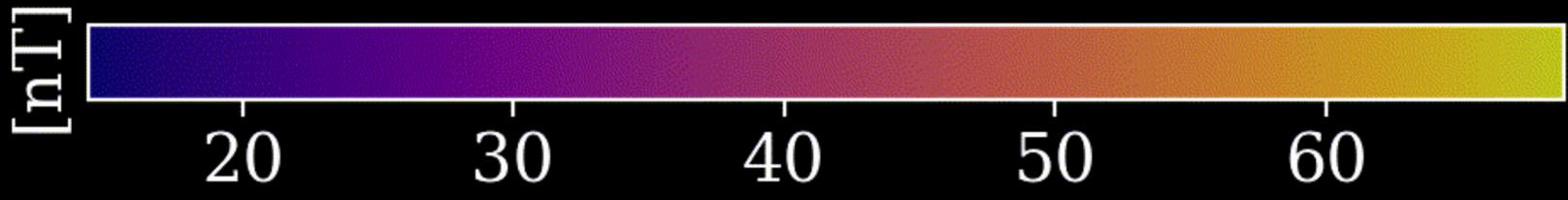
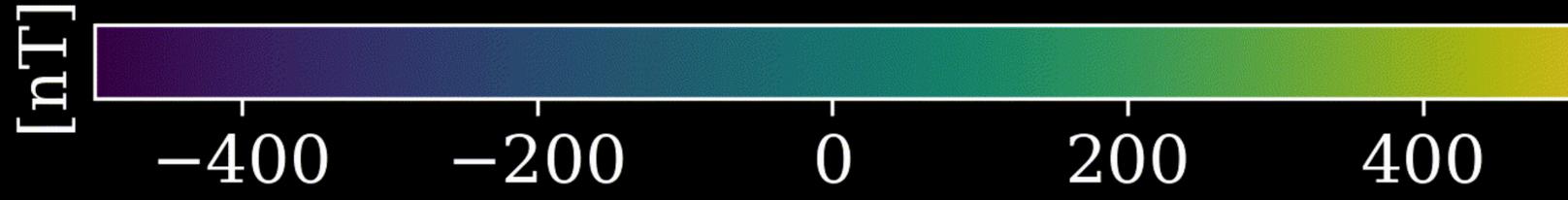
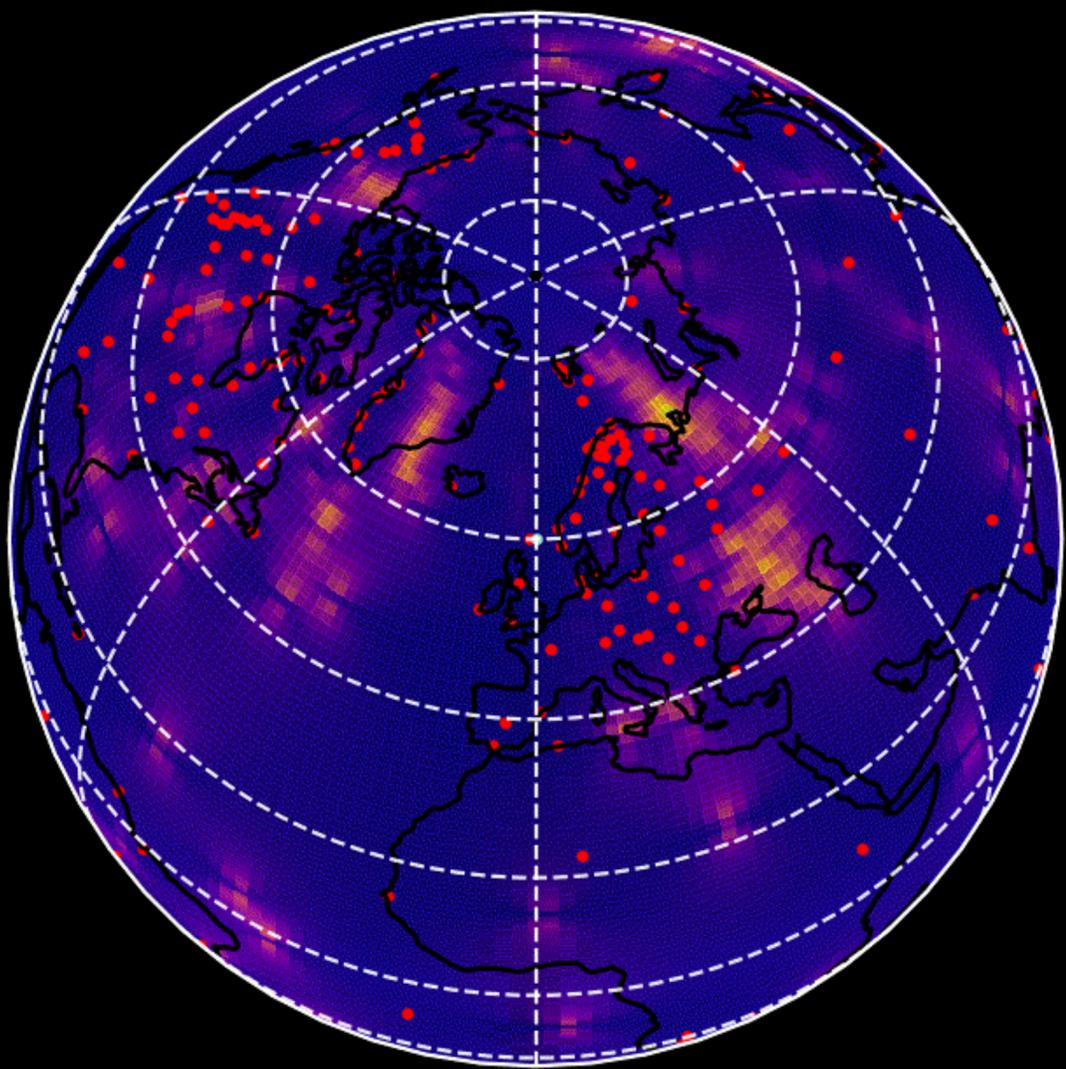


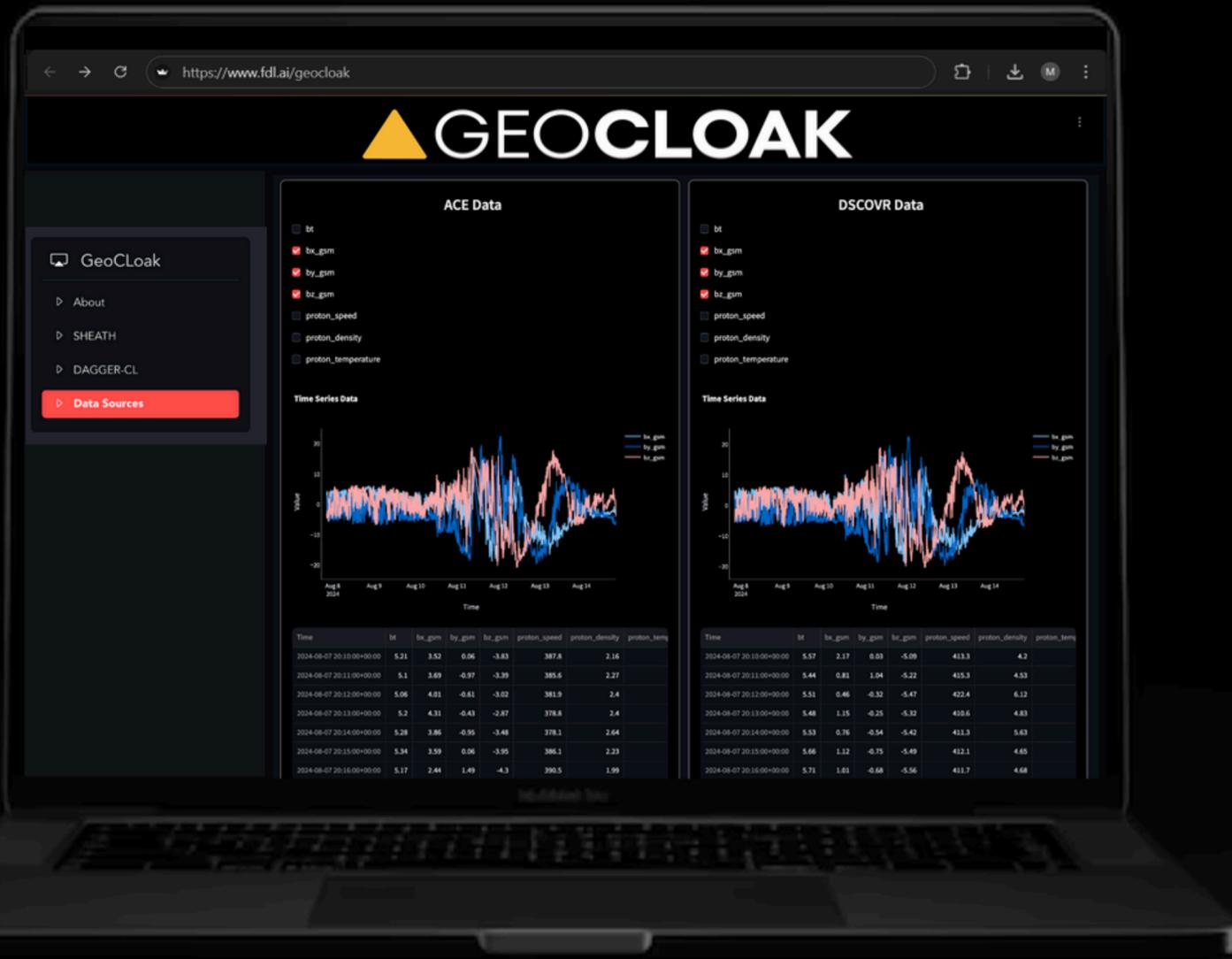
Global Predictions with Quantified Uncertainties

mean B_n



standard deviation B_n

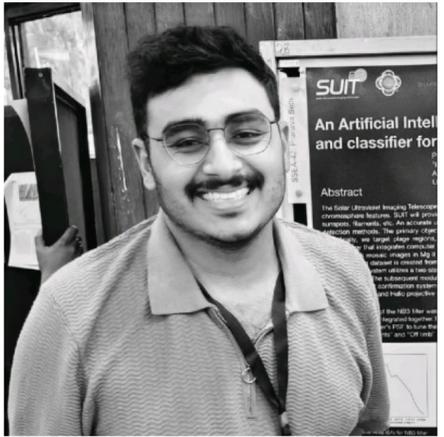
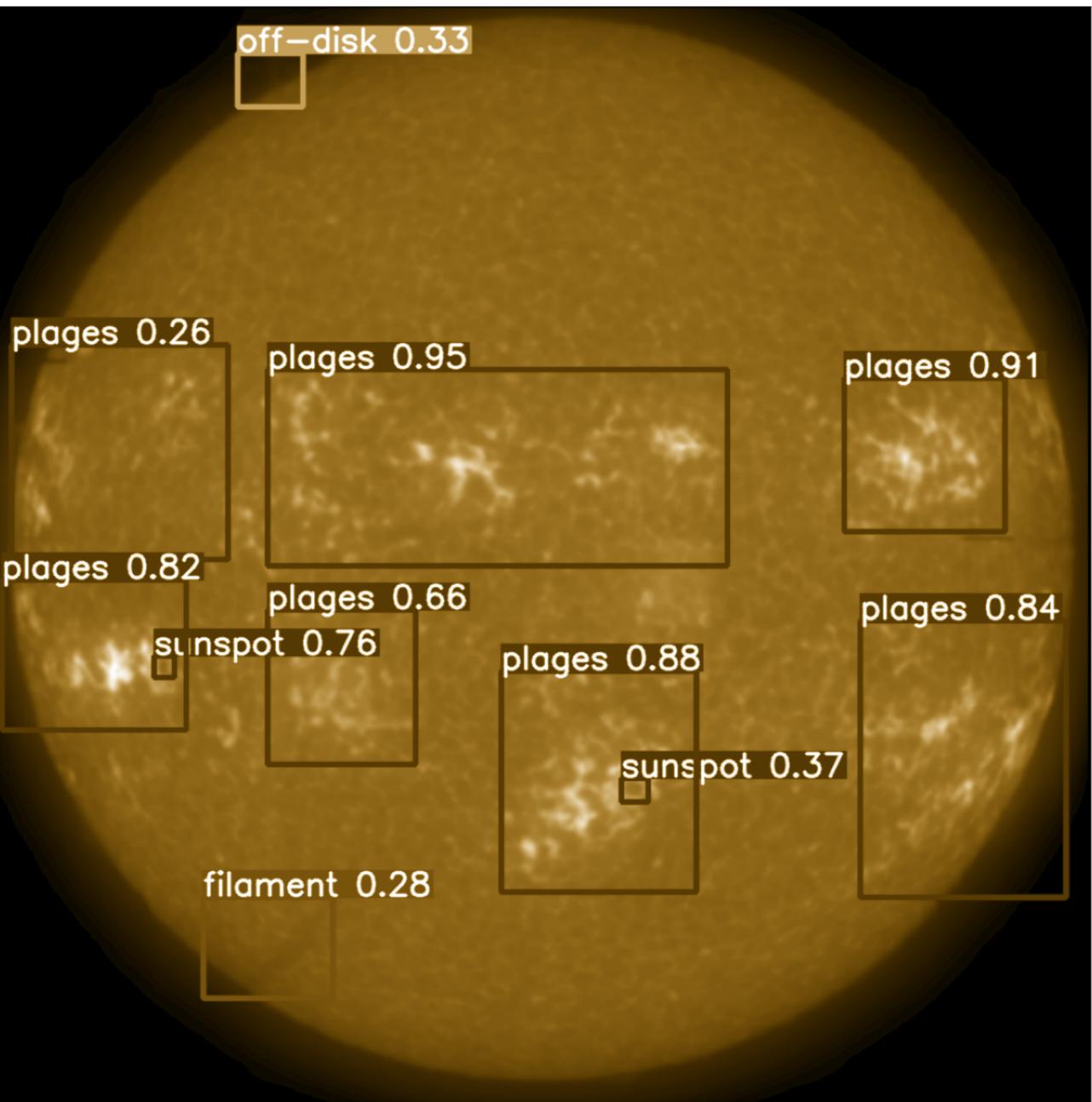
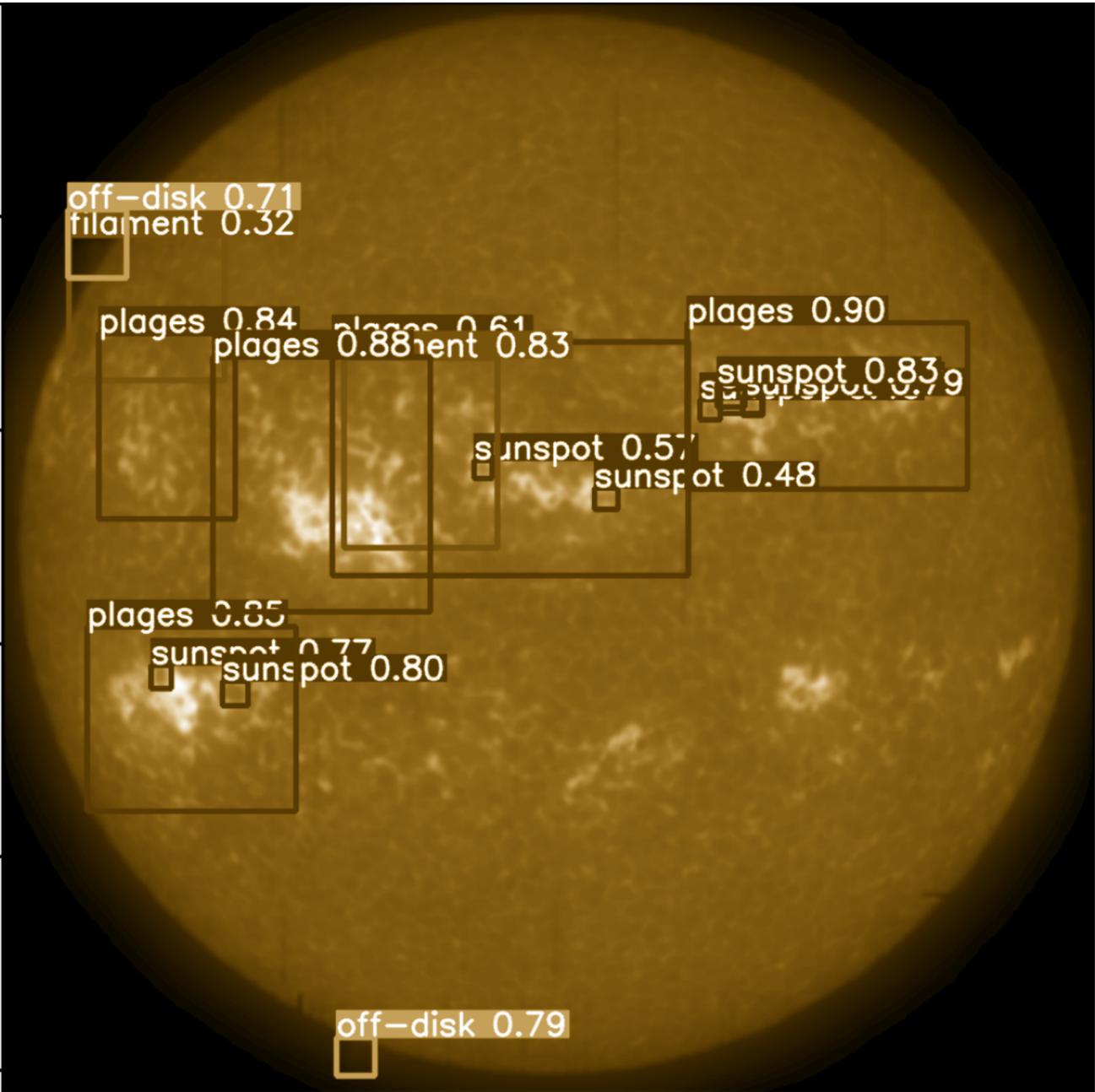




GeoCLOak provides:

- Actionable forecasts of global ground geomagnetic perturbations
- Near real time online inference
- A fit for the future **continual learning (CL)** framework

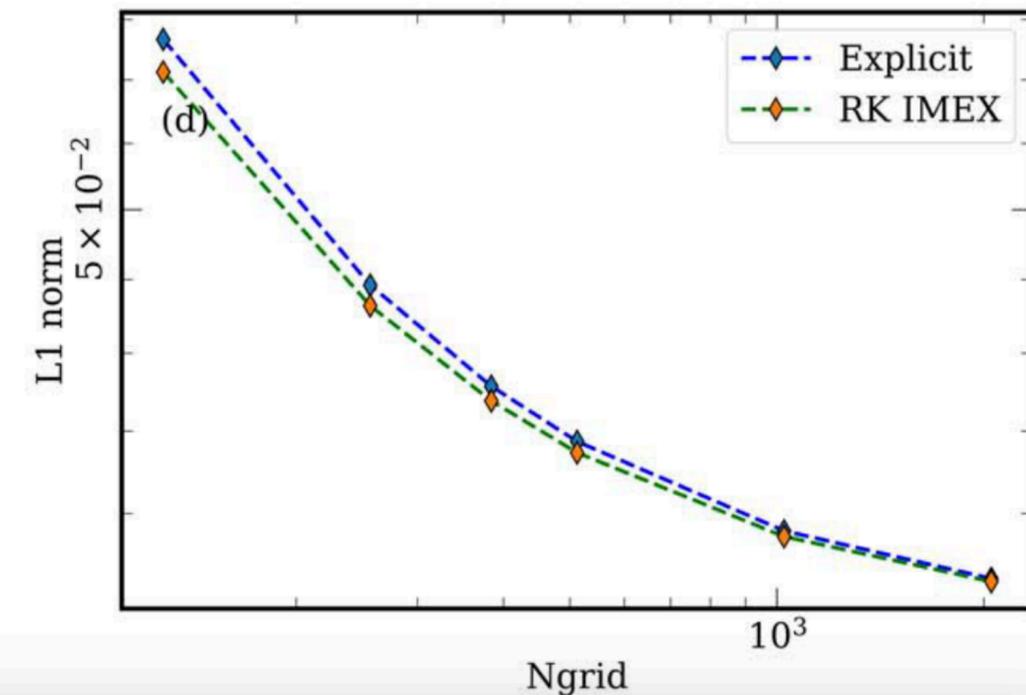
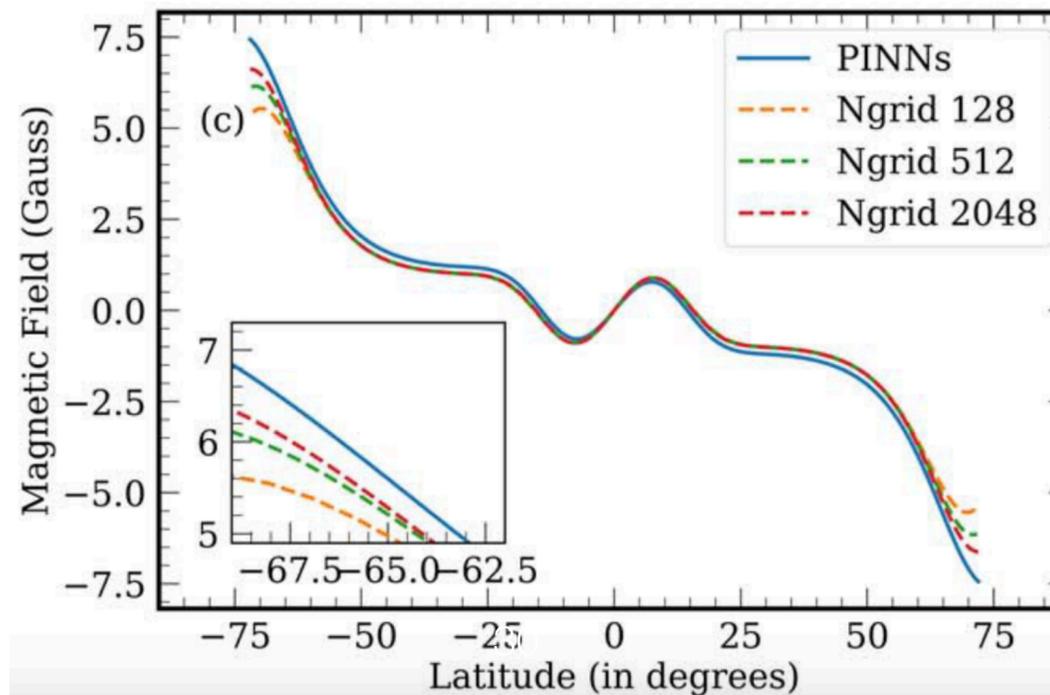
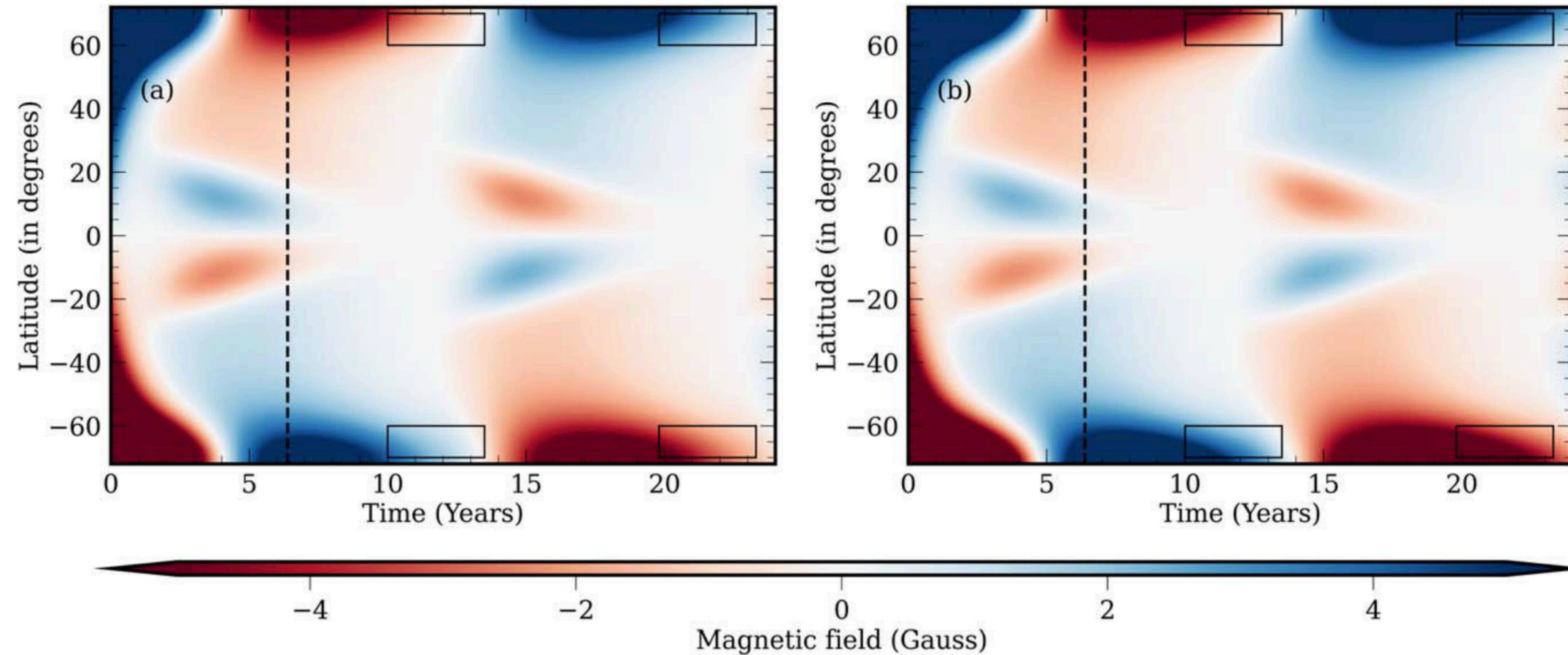
Feature extraction for the Solar Ultraviolet Imaging Telescope (SUIT) onboard Aditya-11



Pranava Seth, Vishal Upendran et. al SolPhys (under review): **The SUIT Team**

Physics-Informed Neural Networks for advection-diffusion equation

- Classical solver tends to PINN solution with grid resolution.
- Diffusive nature of classical solvers - underestimation of polar flux \rightarrow underestimation of solar wind!
- Shows strong potential for data assimilation!



Jithu J Athalathil, ... ,
Vishal Upendran, ..
ApJ 2024 (975:258)

Key takeaways

- Machine/Deep learning can be potentially used to **accelerate inference, provide very fast predictions and at very high cadence** - accelerated by large datasets and compute.
 - Can either be used for simulation based inference or for generating large labelled datasets.
- **Deep models learn some aspects of the physics of the system:** Needs careful quantification, evaluation and validation through physics.
 - **Solar wind sources** are clearly seen, while **solar flare triggers** are potentially estimated.
- Alternatively, formulate model scheme by **baking in physical constraints** to generate realistic constraints on parameter space.
- Cross disciplinary collaboration and open source data / open science methodology critical !

Paths to future

- **Simulation based inference:** Inference using physics based statistical models to constrain properties of small-scale impulsive events [Rajhans et al, ApJ, under review].
- **Interpretable AI for solar wind sources:** Constrain source regions using solar wind composition and multiple wavelength bands. [Soon!]
- **Interpretable AI for flares:** Temporal coincidence and analysis with magnetic field topology.
- **Physics-incorporated forecasting:** For operations: Bias reduction, change of basis functions, and full evaluation of the continuous learning framework [Kirkpatrick+ 2016, Raghavan+ 2024].